

Planck's formula

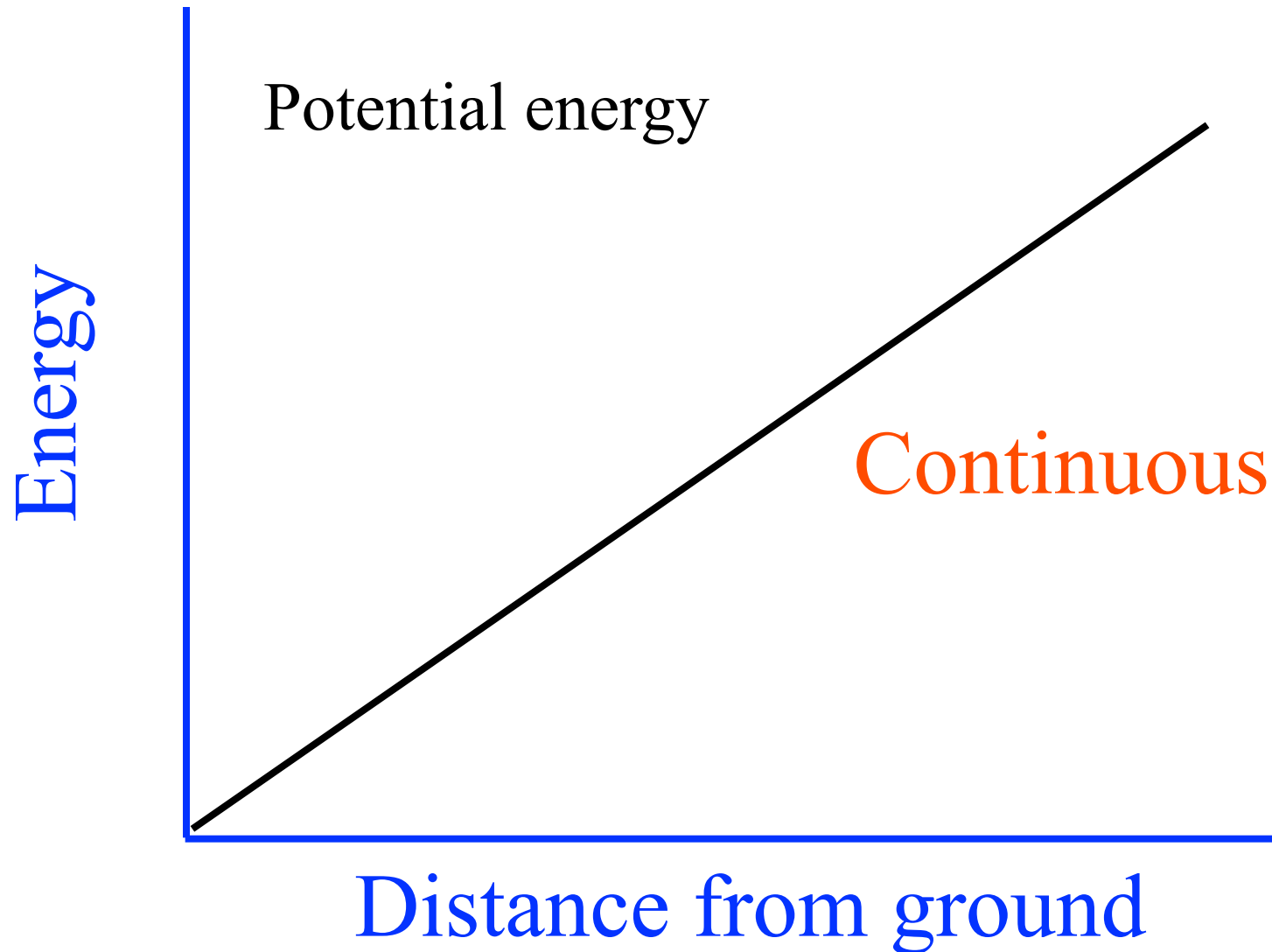
In 1900 Planck derived a formula that matched the measured curve exactly.

His derivation was based on the blackbody having energy states that are *quantized*.

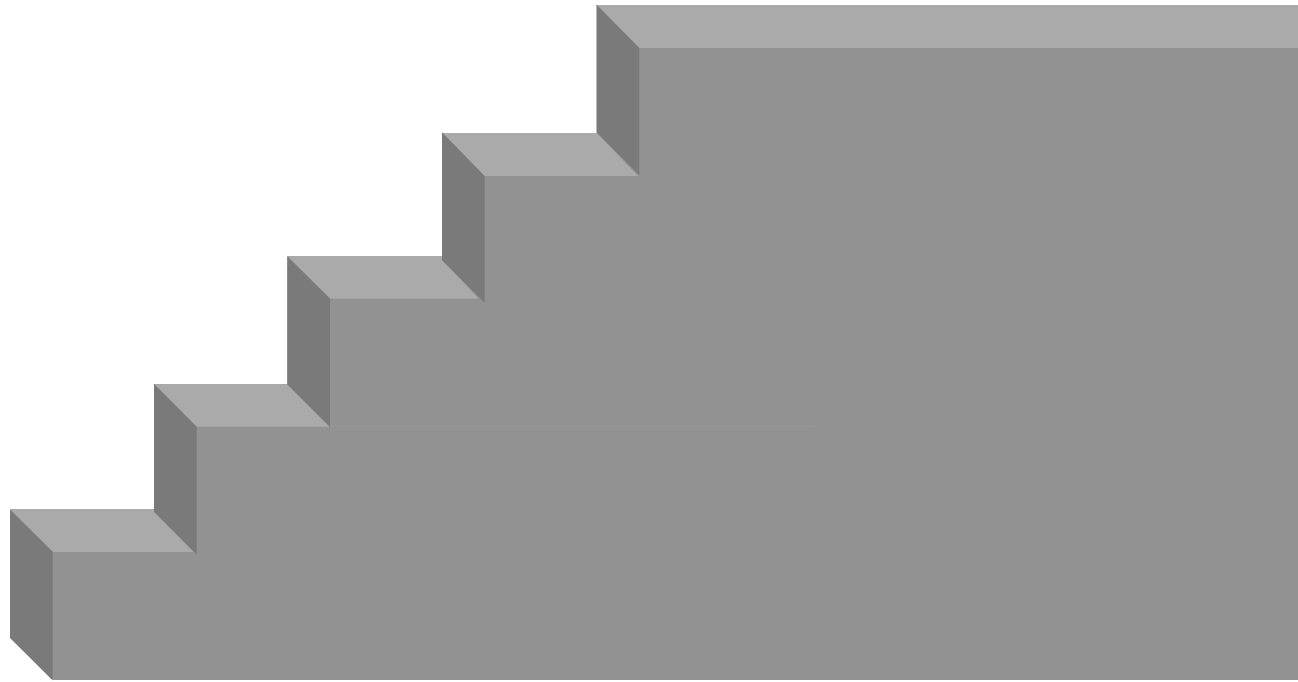
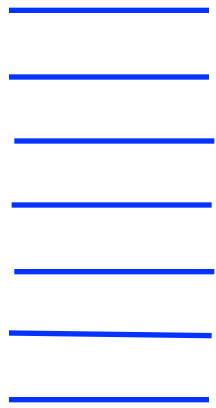
$$E = n h \nu$$

$$n = 1, 2, 3, \dots$$

Energy in classical physics



Planck's proposed energy



quantized

Spectra of atoms and molecules

Atoms and molecules interact with electromagnetic radiation.

Energy of the radiation at certain frequencies can be absorbed when it hits the atom or molecule

Energy is conserved so it has to go somewhere

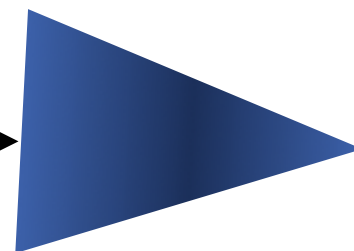


source

I_0 →

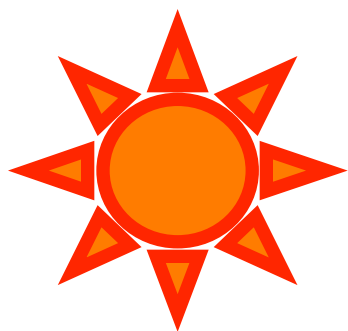
no gas

I_0 →



detector





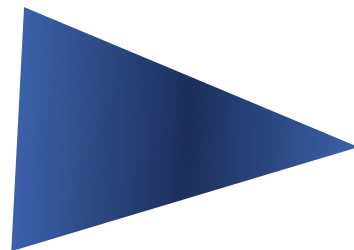
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I_0



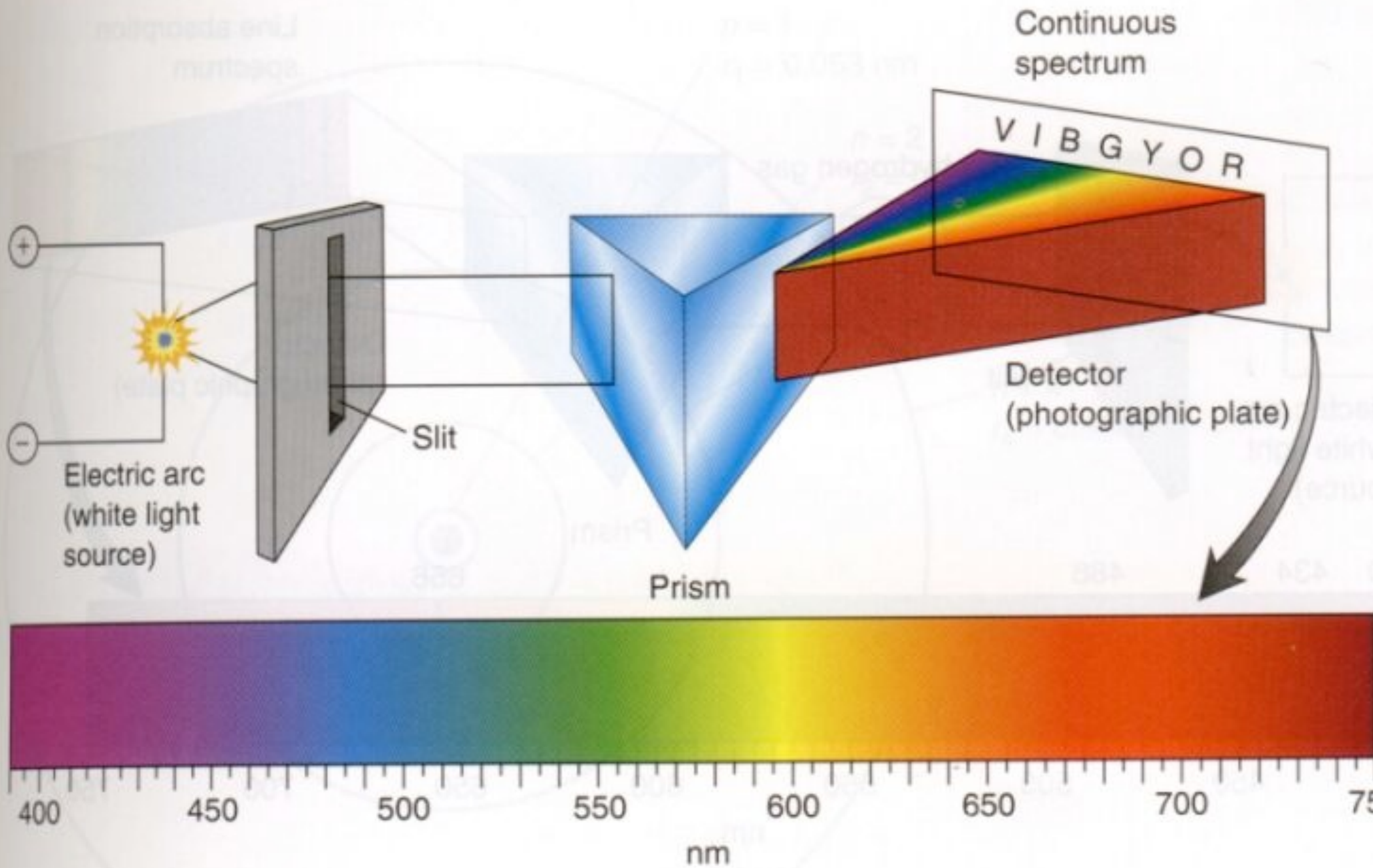
gas

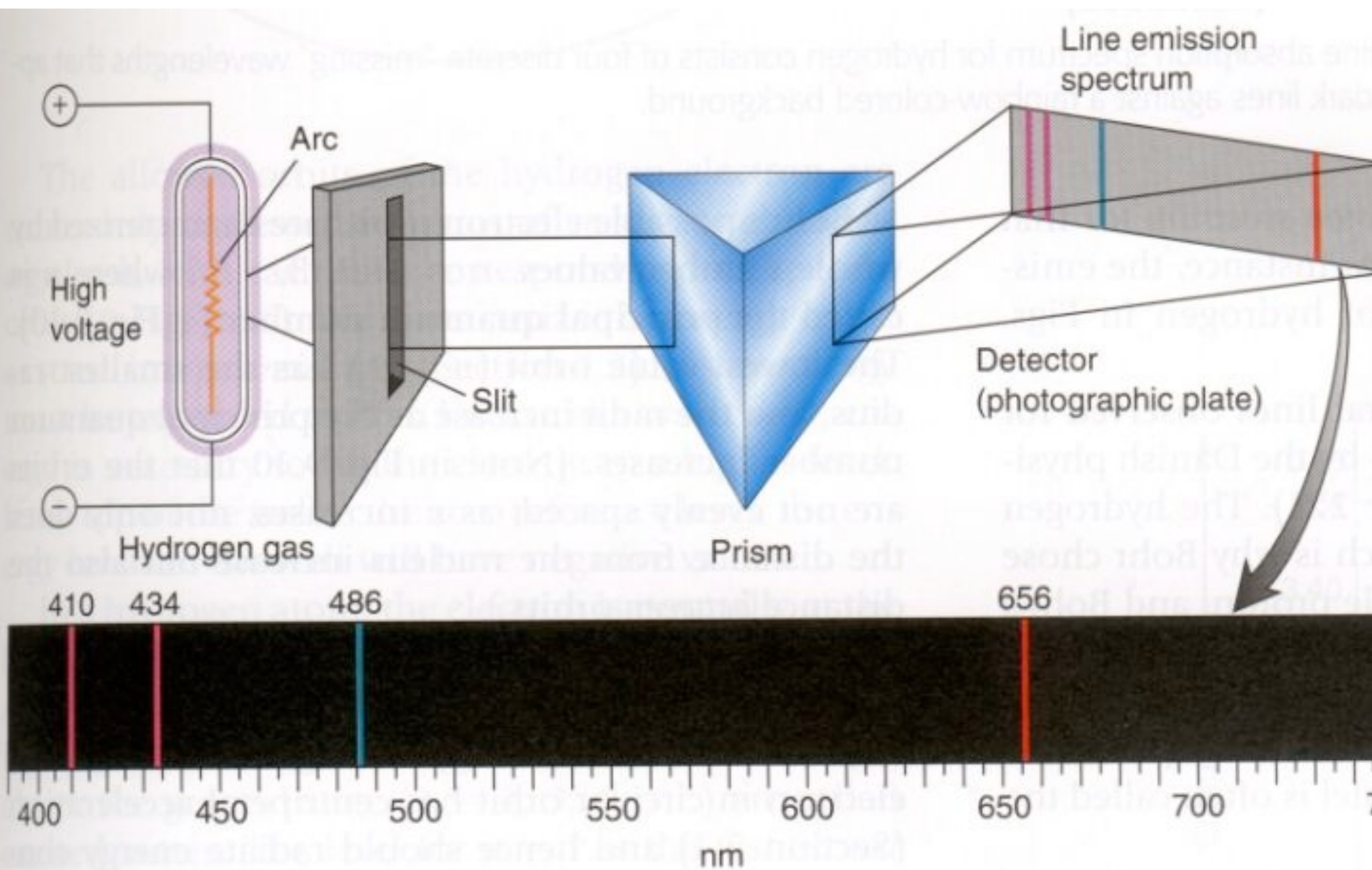
$I < I_0$

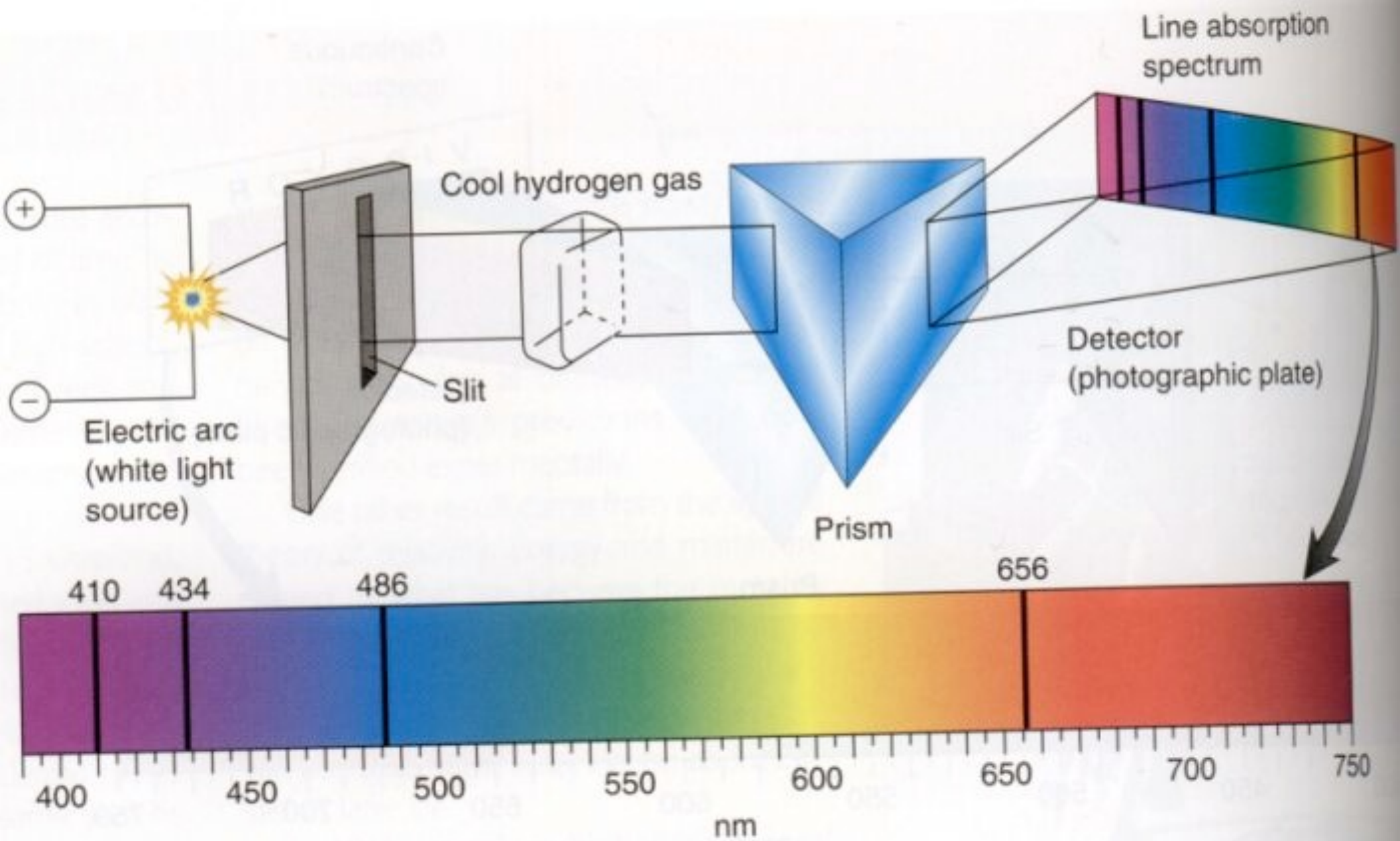


detector







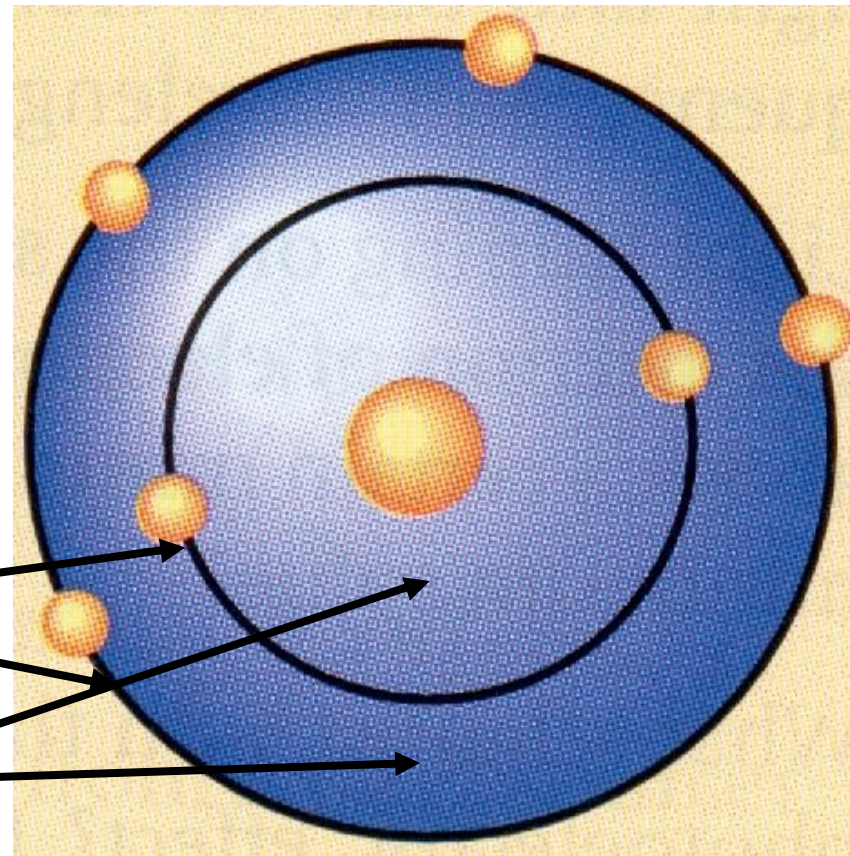


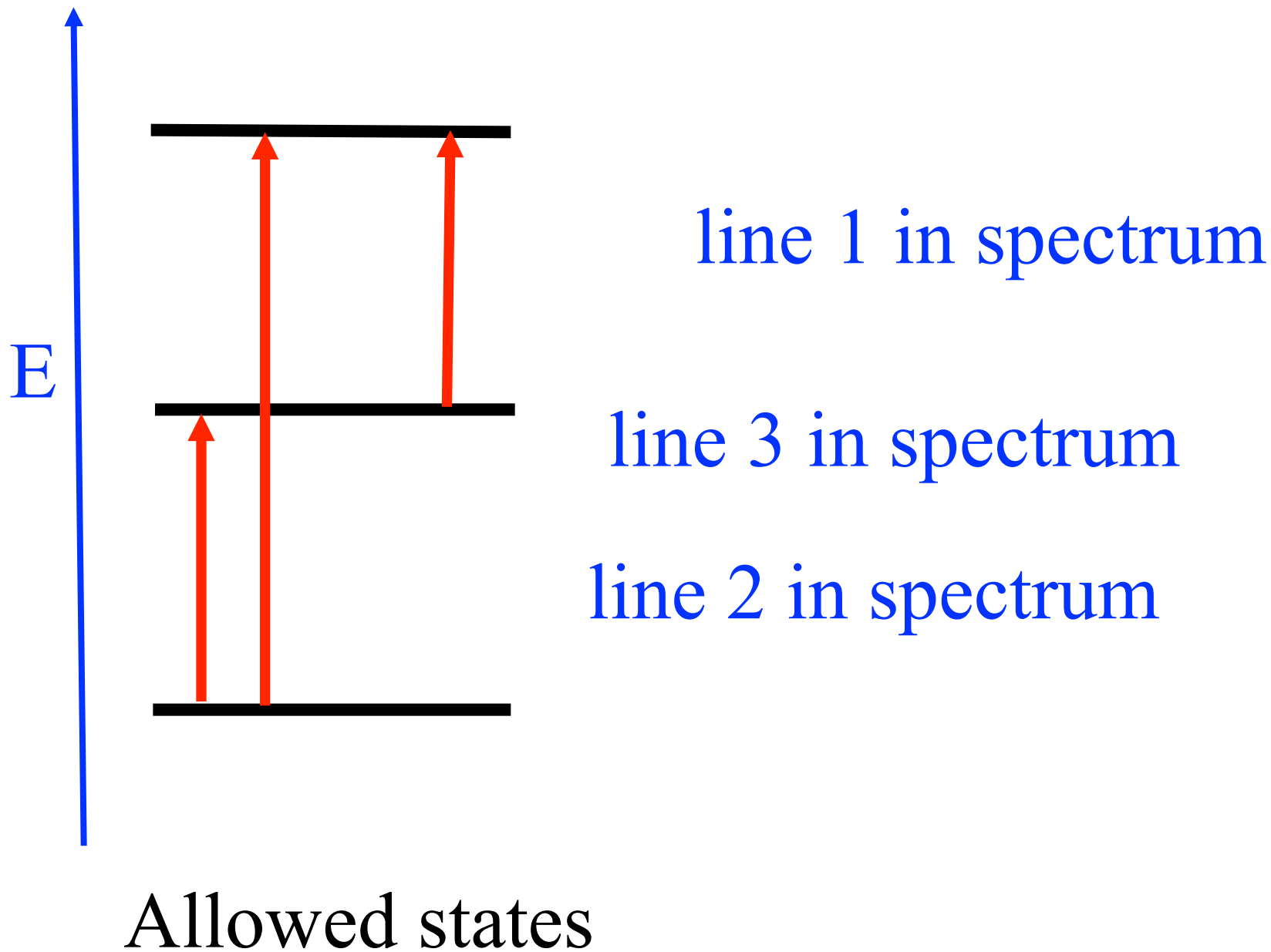
Explanation: quantization of energy of atoms and molecules

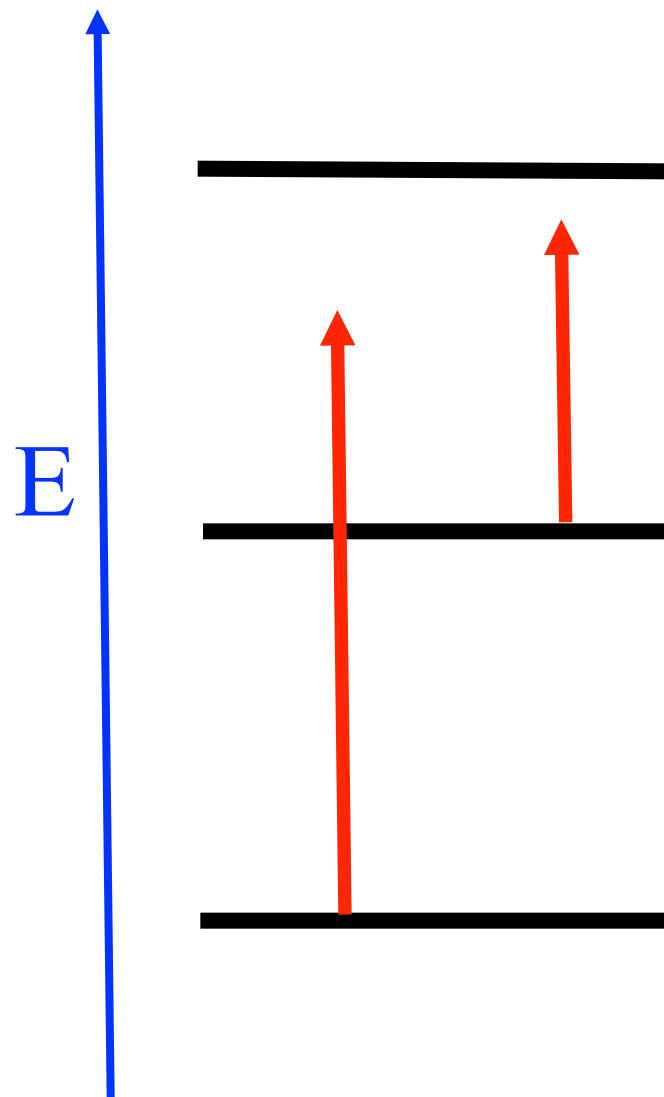
First model was Bohr's planetary model of the atom

Allowed energy state

Disallowed energy state







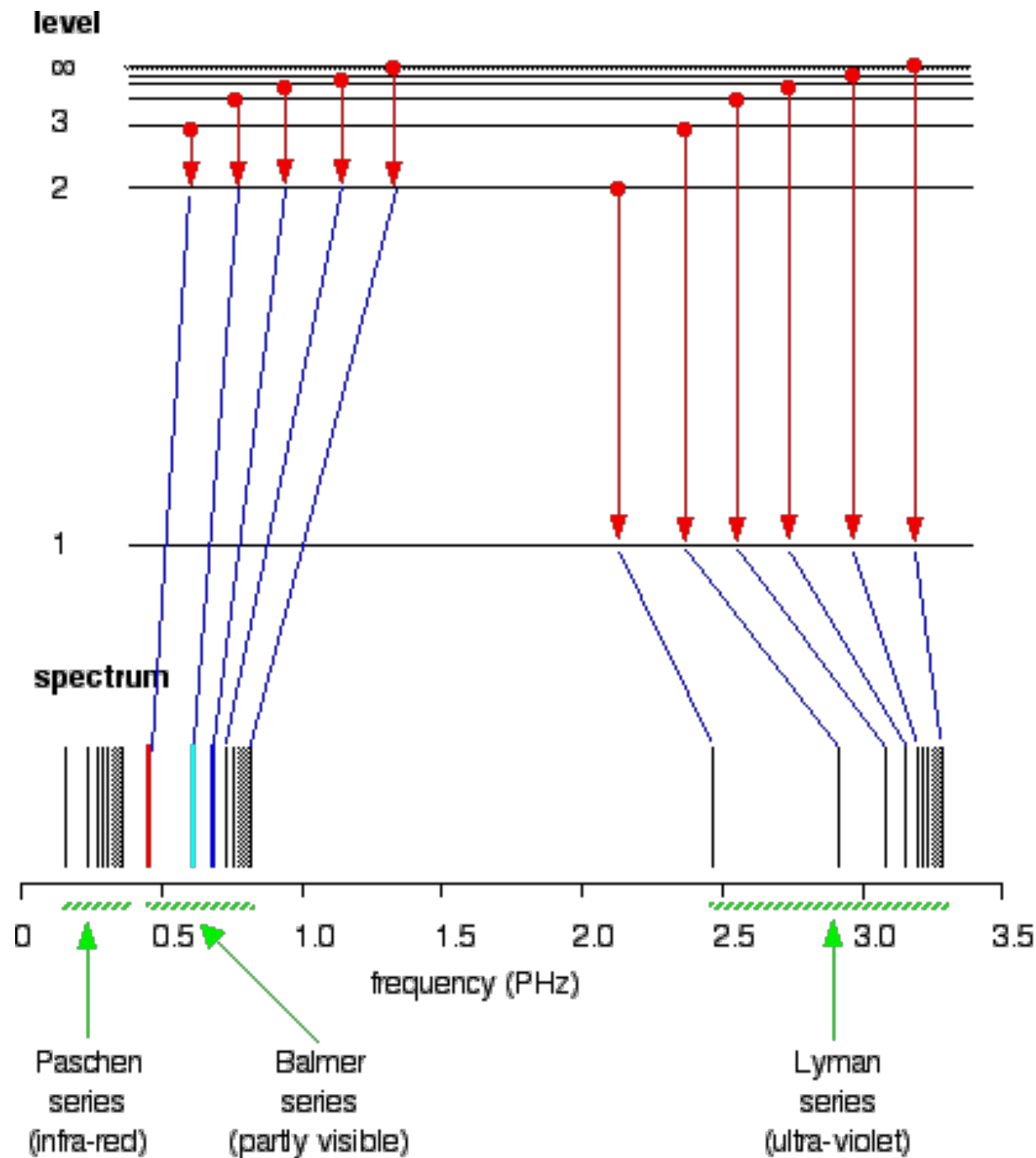
line 1 not observed

$h\nu$ of the photon must be very close to the energy of separation of the levels for a transition to occur.

Allowed states

What was observed - Electronic Spectra

Electronic Spextra



Electronic Spectra

$$\frac{1}{\lambda} = R \left(\frac{1}{(n'')^2} - \frac{1}{(n')^2} \right)$$

where R is the Rydberg constant:

$$R = 1.097373 \times 10^7 \text{ m}^{-1}$$

Microwave – molecular rotation

Electronic Spectra

when $n''=1 \Rightarrow$ Lyman series

when $n''=2 \Rightarrow$ Balmer series

when $n''=3 \Rightarrow$ Paschen series

when $n''=4 \Rightarrow$ Brackett series

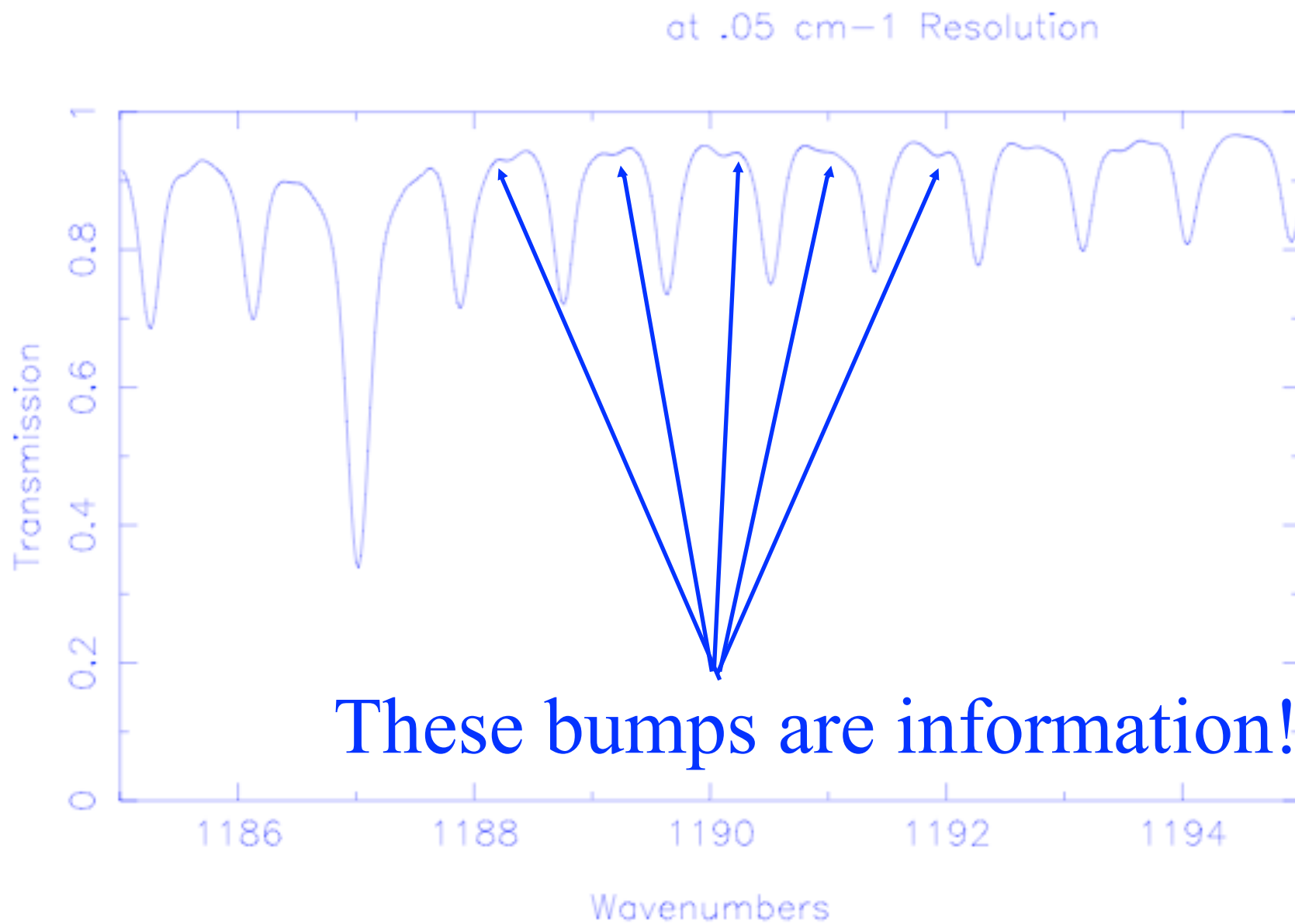
Visible region – electronic spectra

Infrared – molecular vibration and rotation

Microwave – molecular rotation



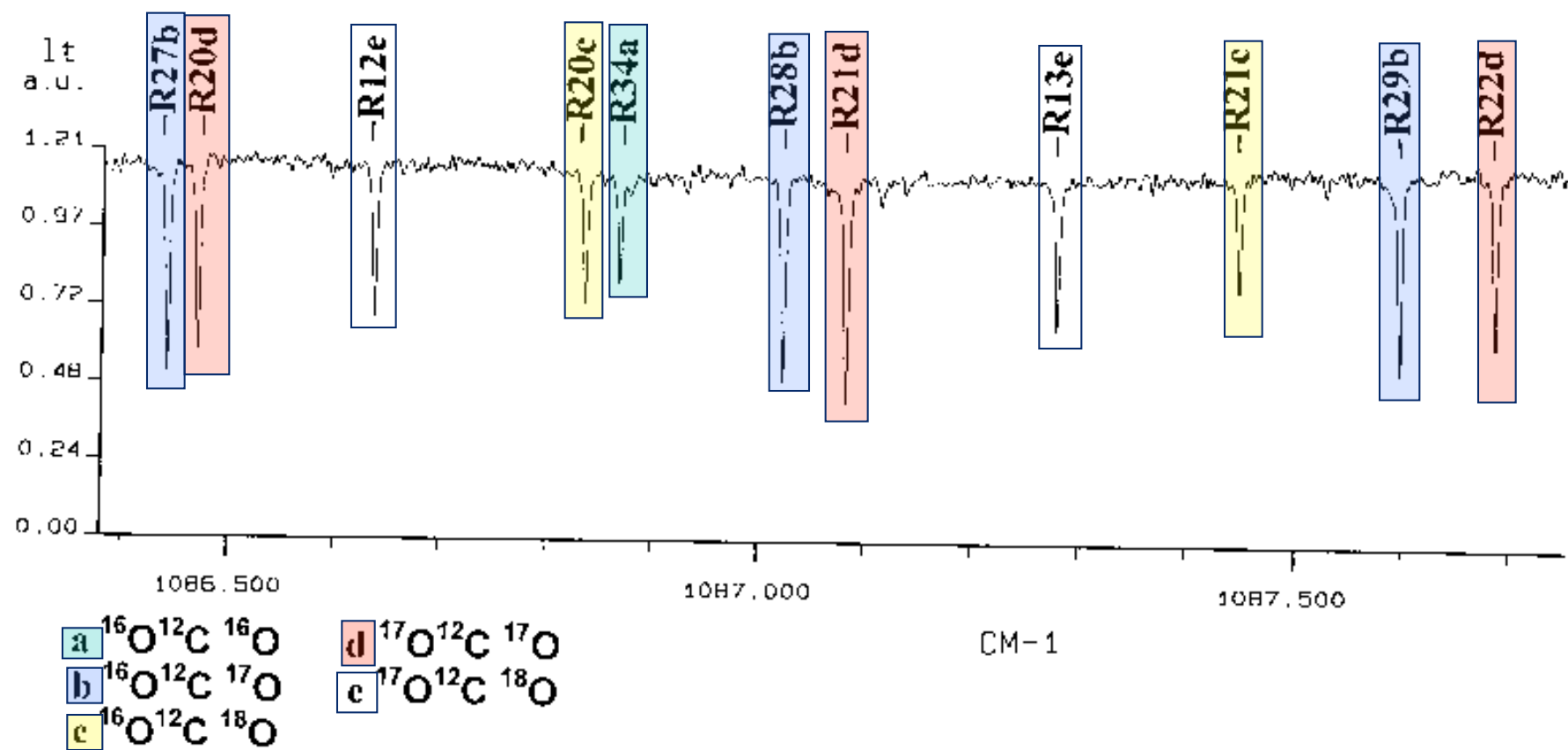
What do we see?



Carbon Dioxide

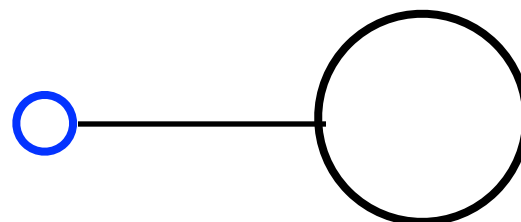
Isotopically enriched sample ^{17}O



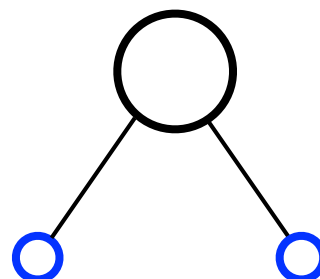


molecules with dipole moments

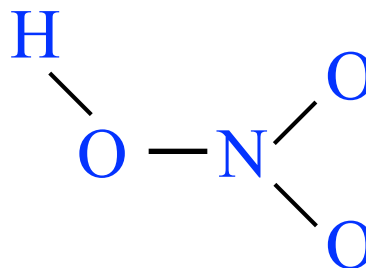
HCl



H₂O



HNO₃



molecules with quadrupole moment
more symmetric, cylindrical distribution

examples:

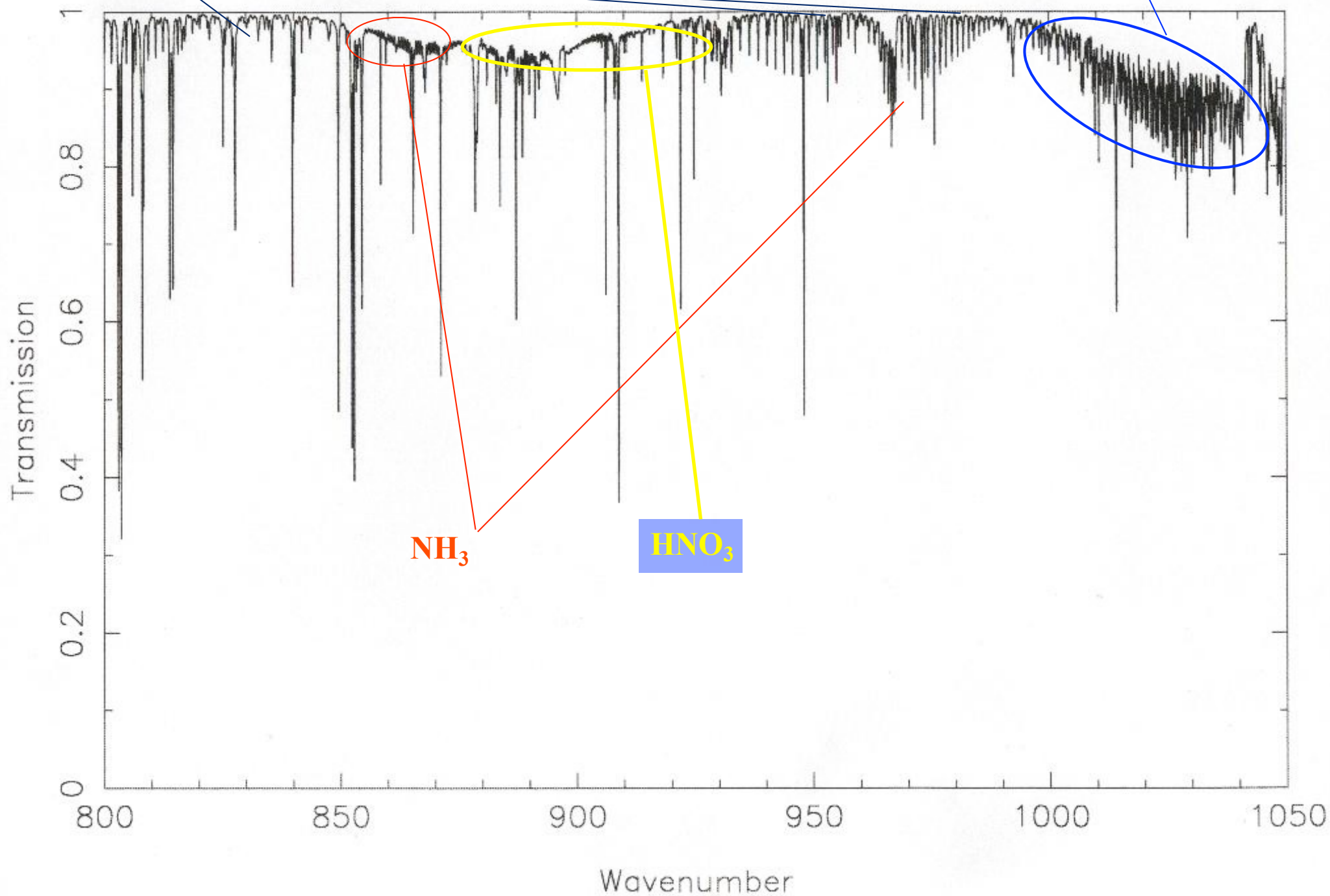
N_2



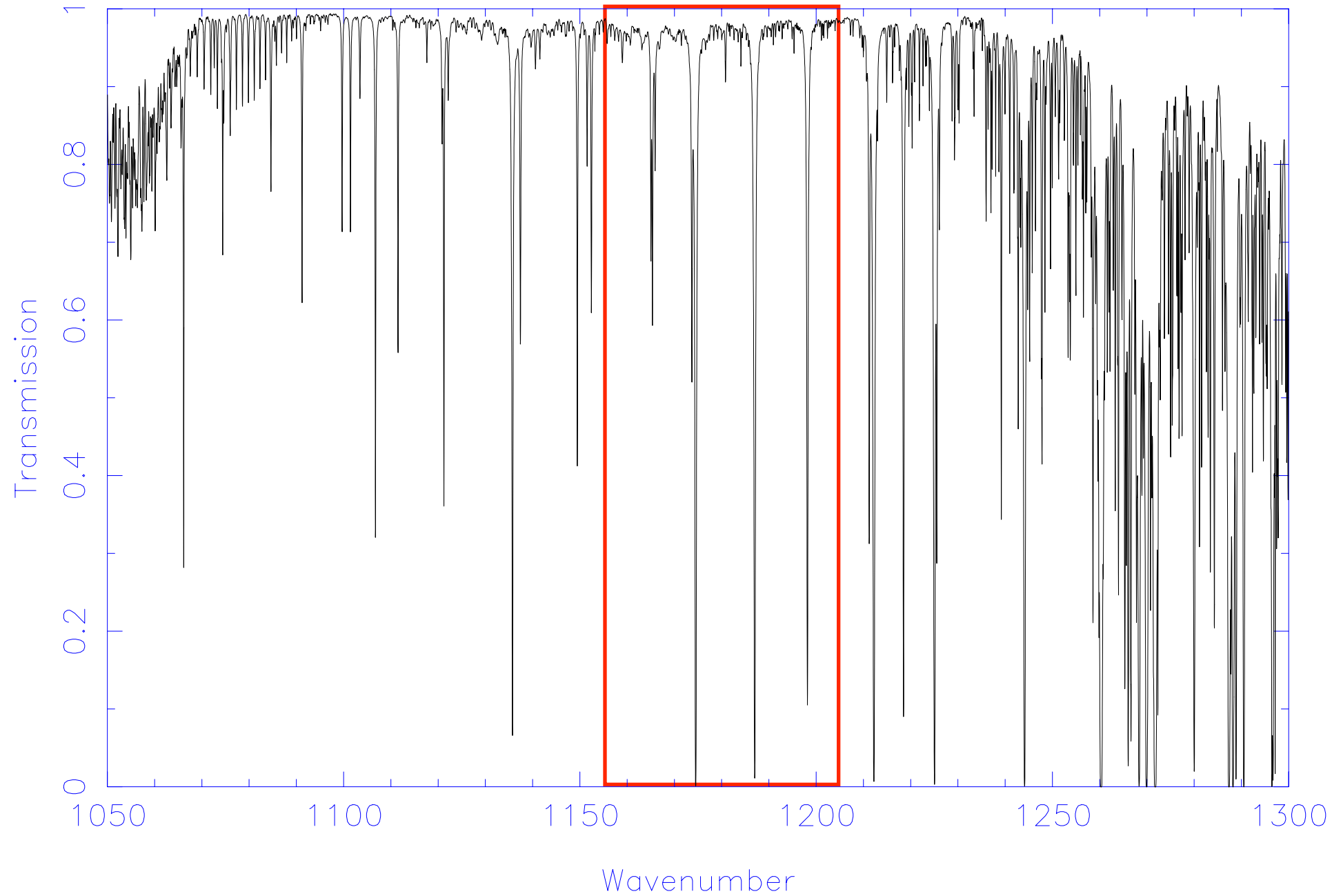
CO₂

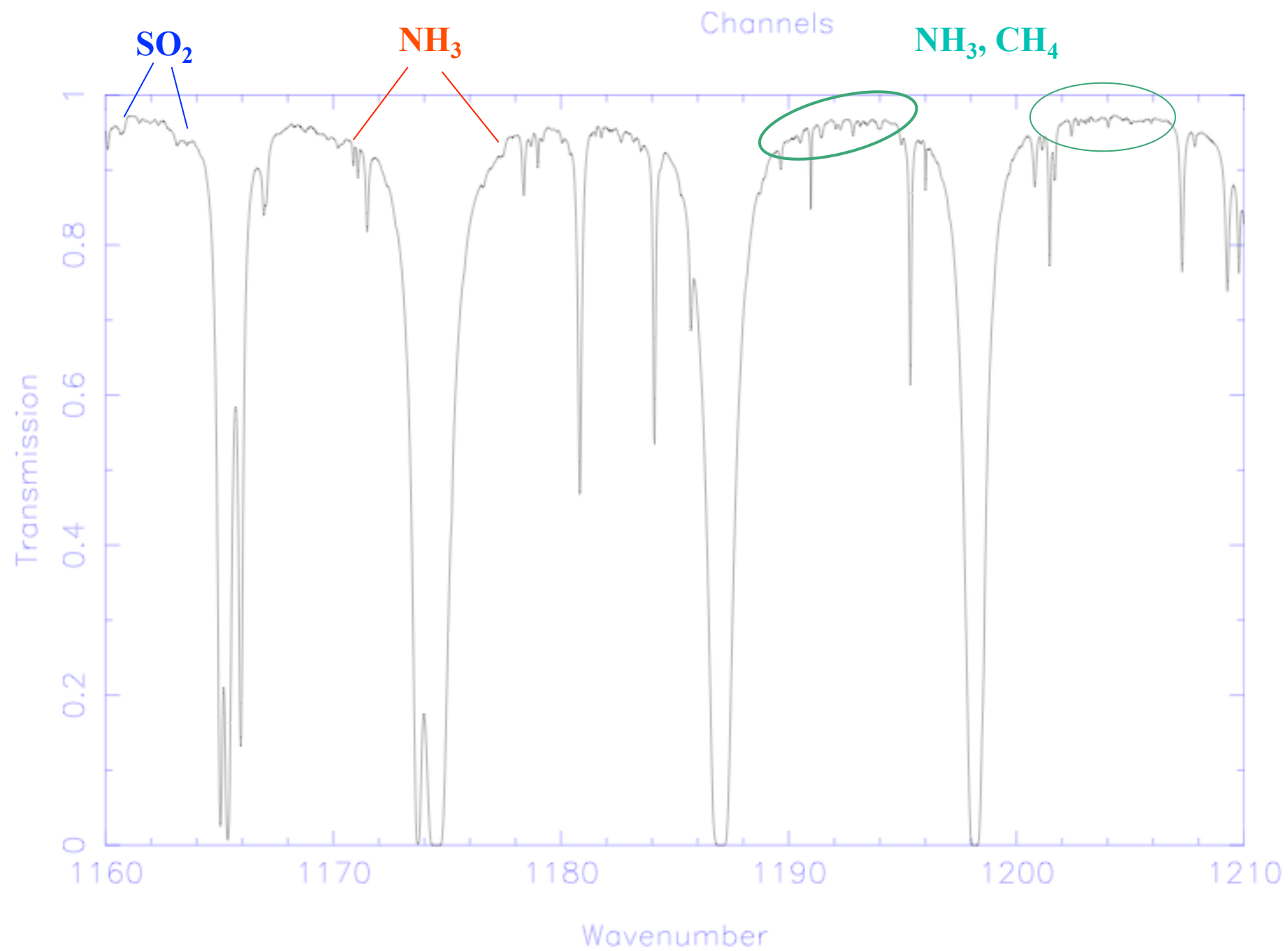
Standard Atmosphere, 1000 RH=05

O₃

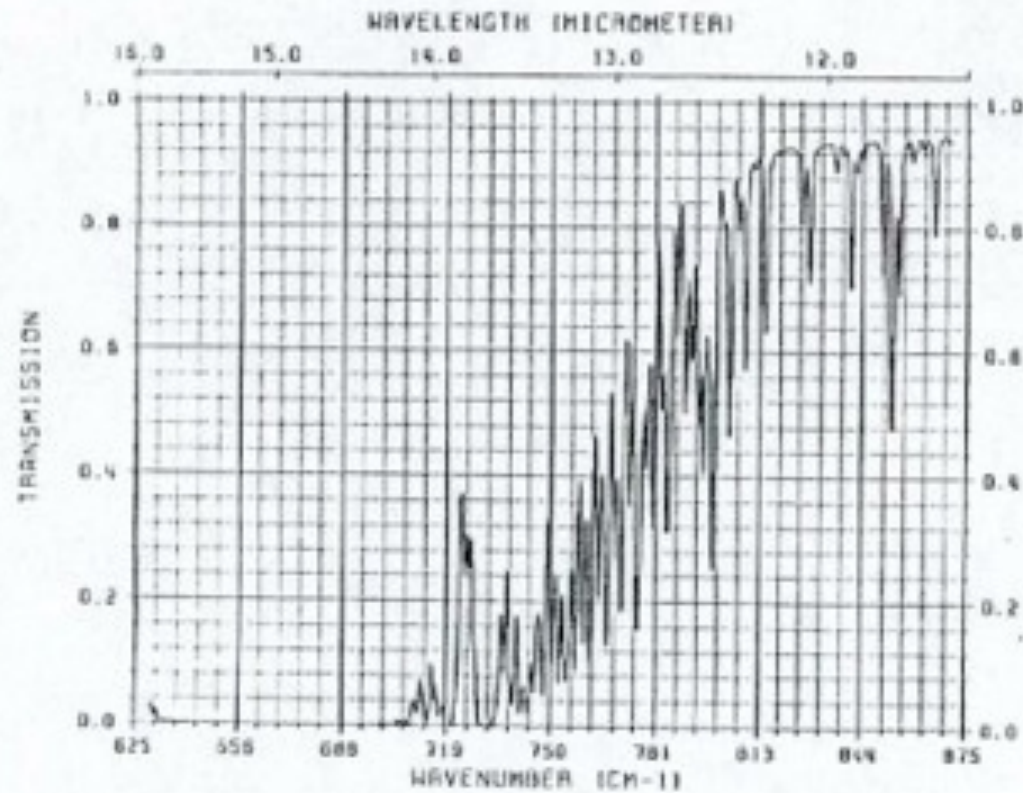
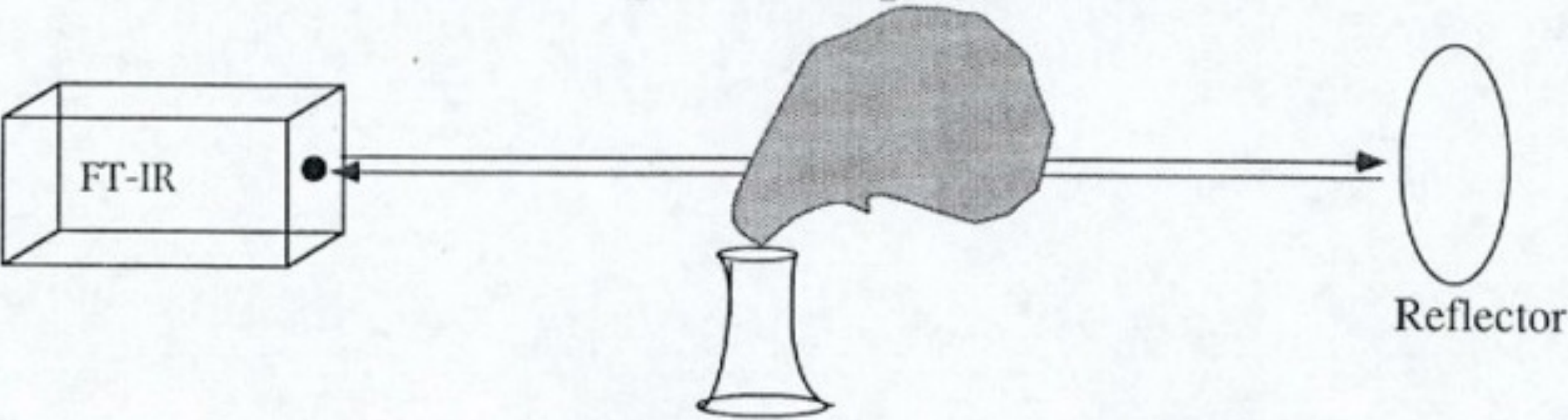


Standard Atmosphere, 1000 RH=05

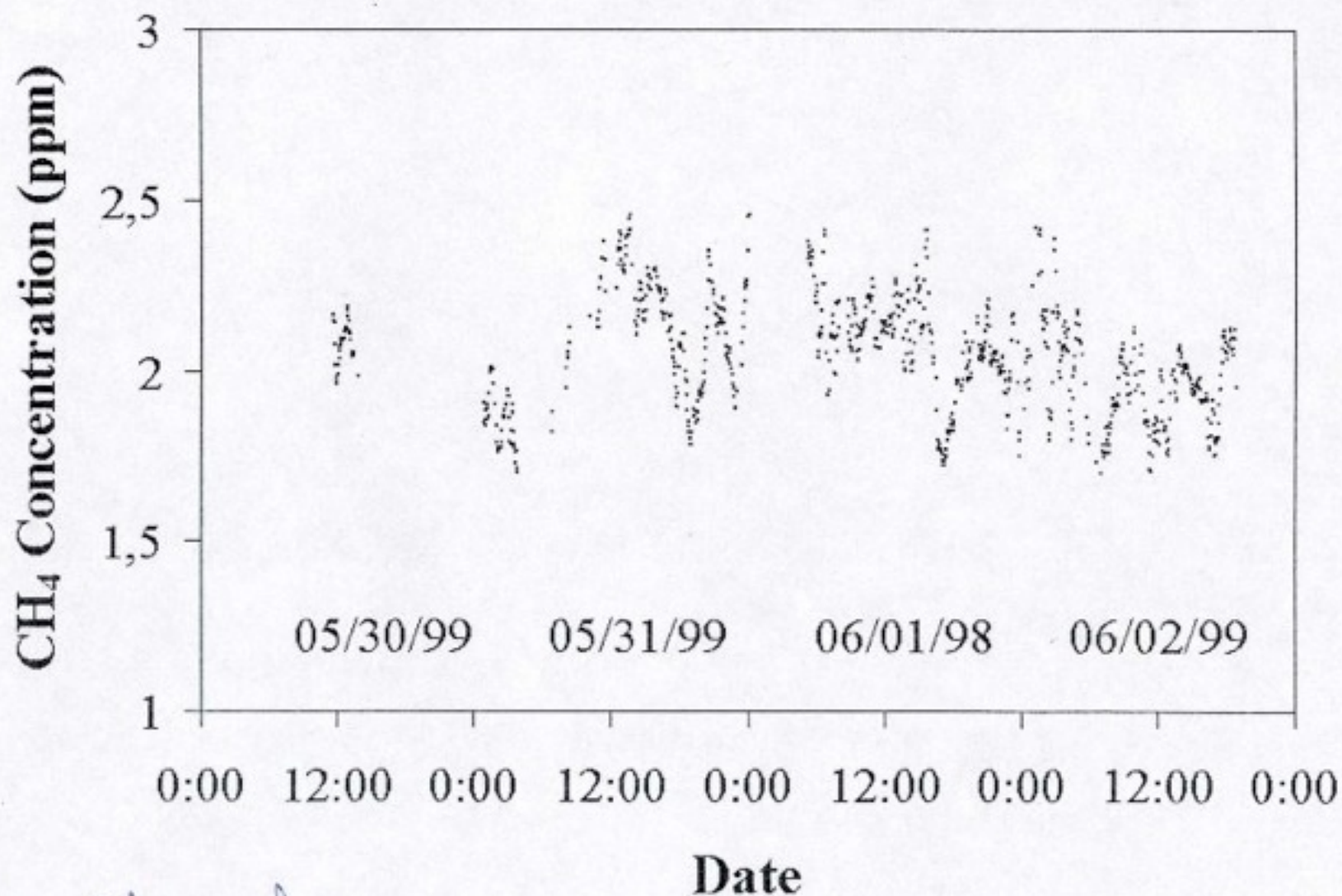


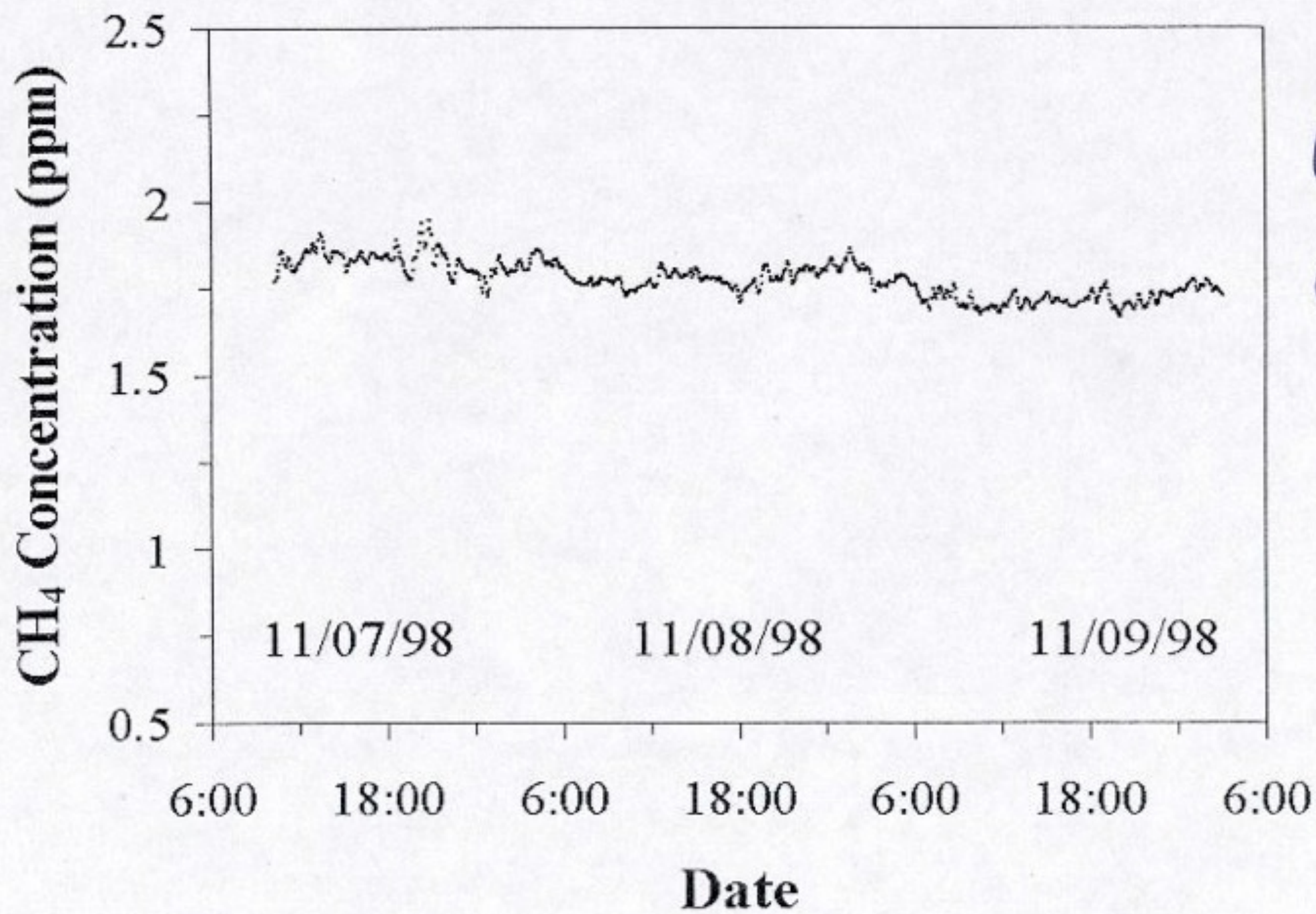


Remote Sensing of Atmospheric Pollutants



“Lac Piché” (June 99)





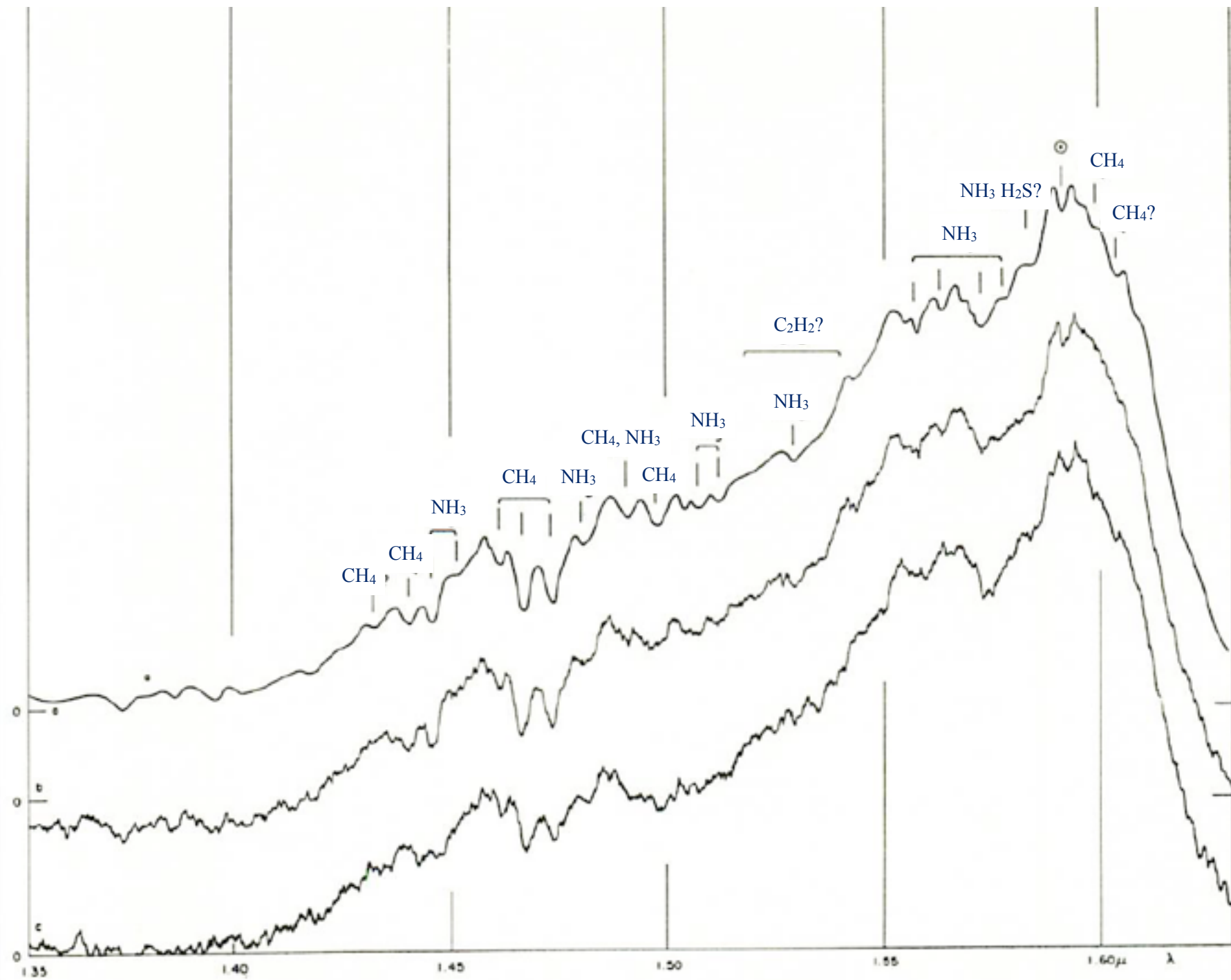
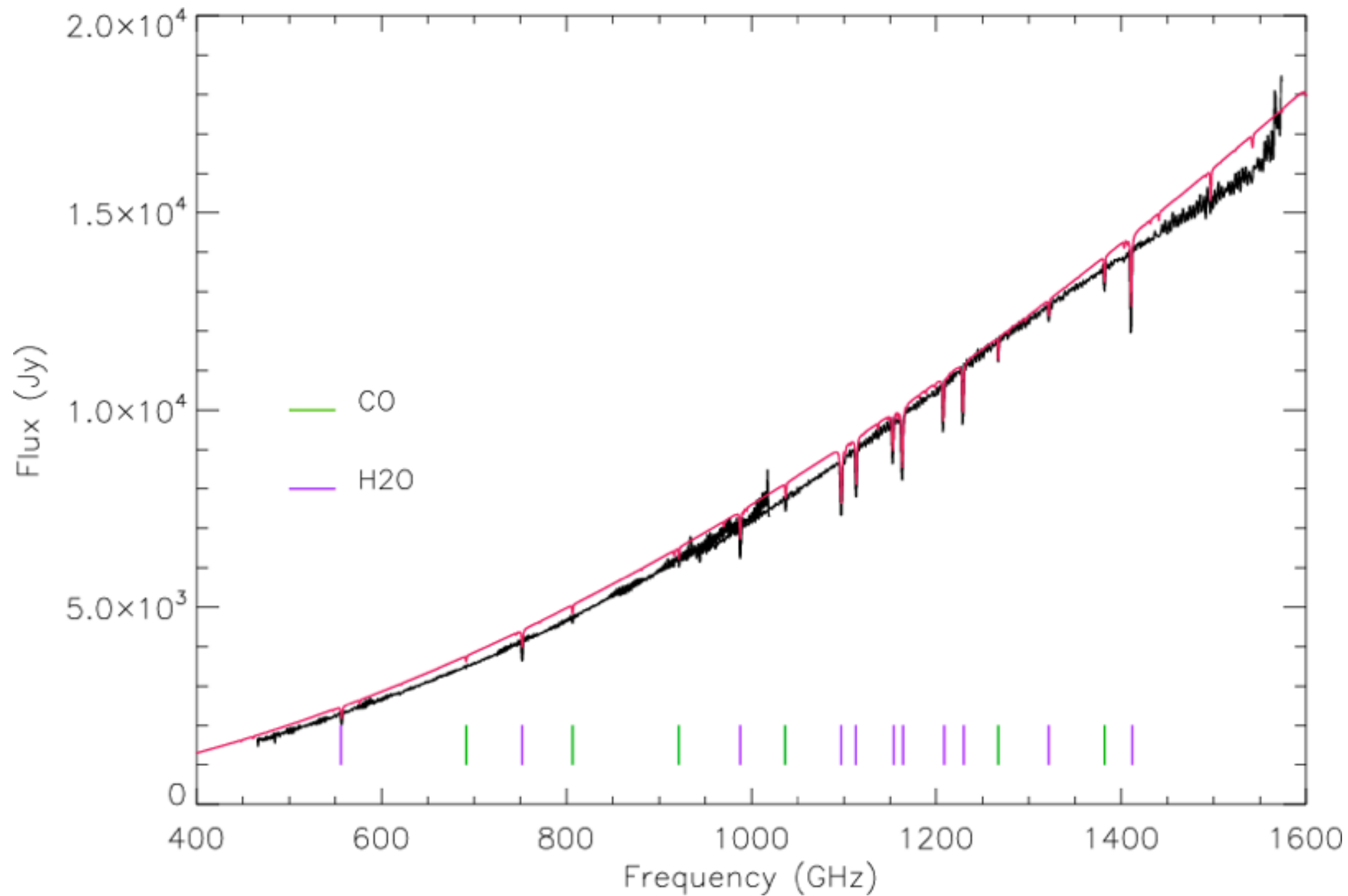
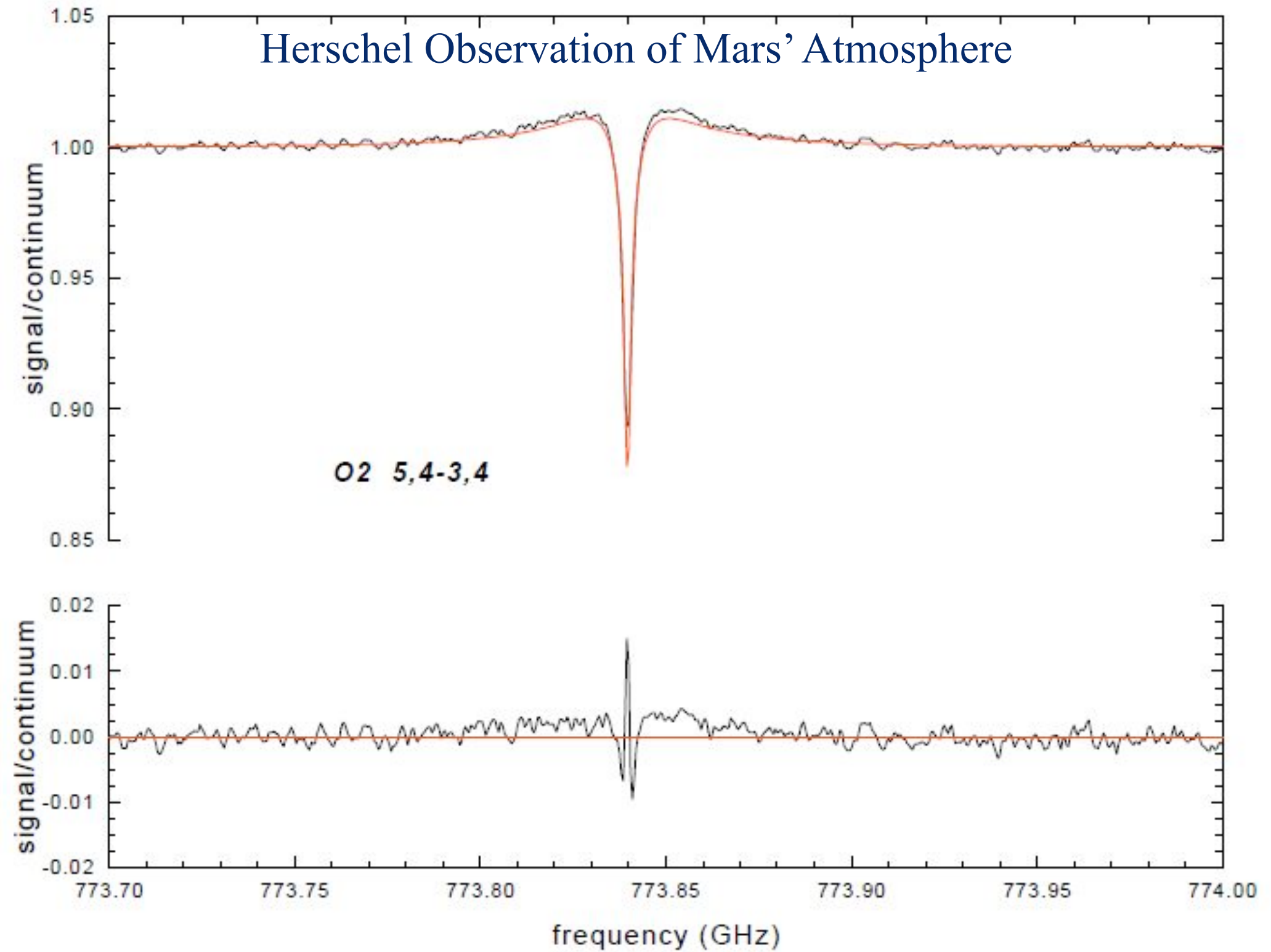


Fig. 3 Spectrum of Jupiter, Part 3, 1.35–1.63 μ ; as Fig. 1.

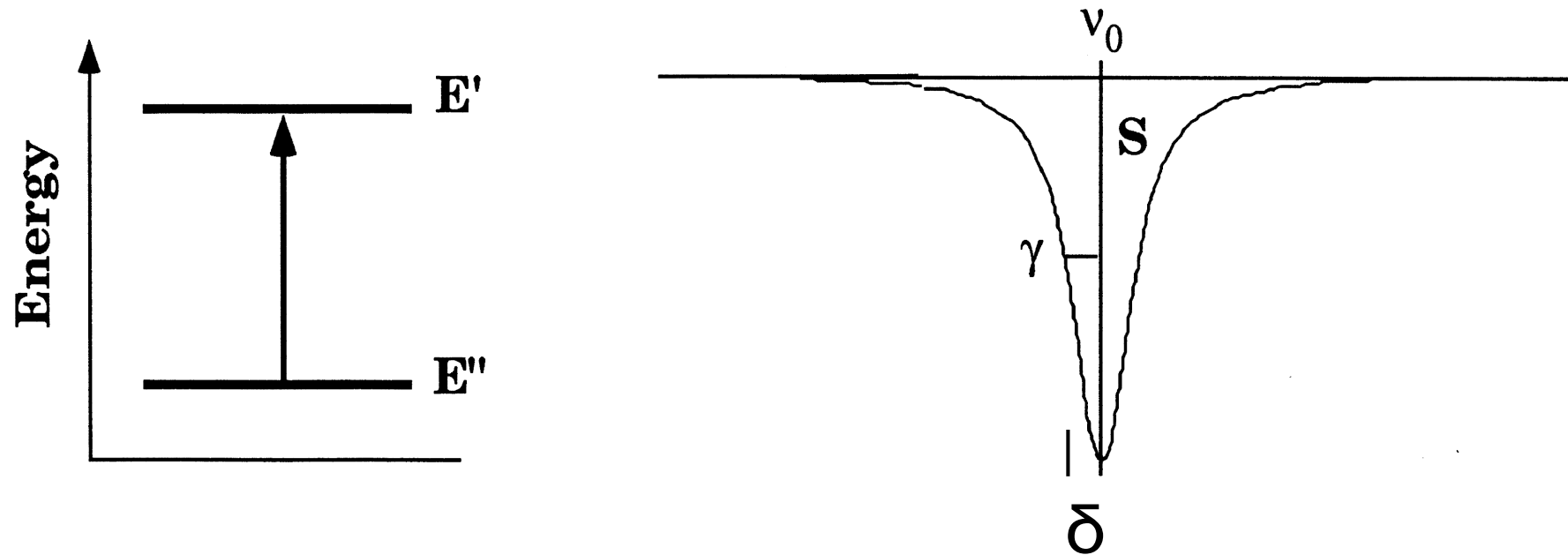
Herschel Observation of Mars' Atmosphere




Herschel Observation of Mars' Atmosphere



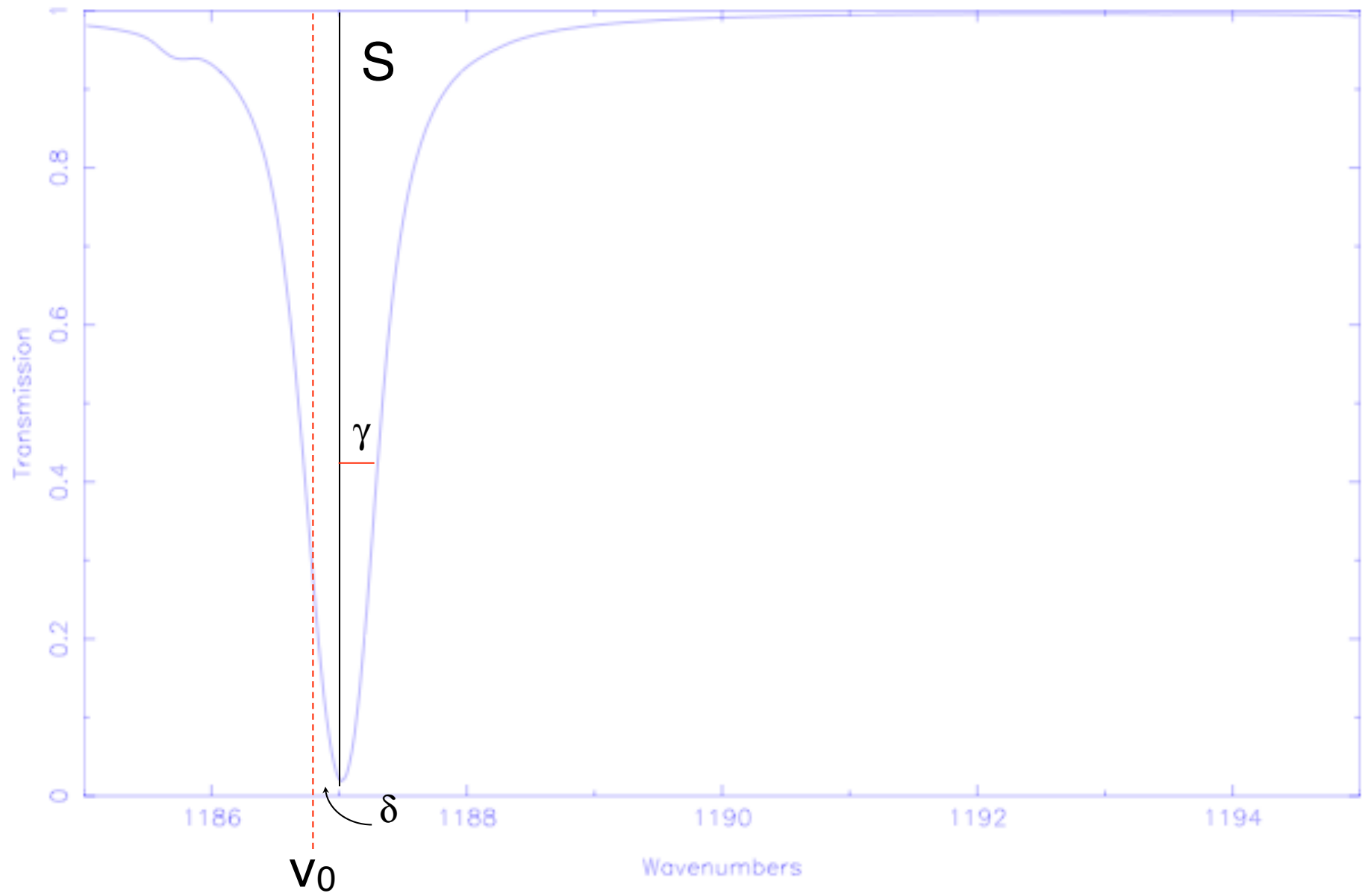
Important Line Parameters



Quantized Energy Transition  Broadened Lorentzian Line Shape

H₂O

at 0.1 cm⁻¹ Resolution



Spectral Parameters

ν_0 - vacuum wavenumber (cm^{-1})

S - line intensity ($\text{cm}^{-1}/(\text{molecule cm}^{-2})$)

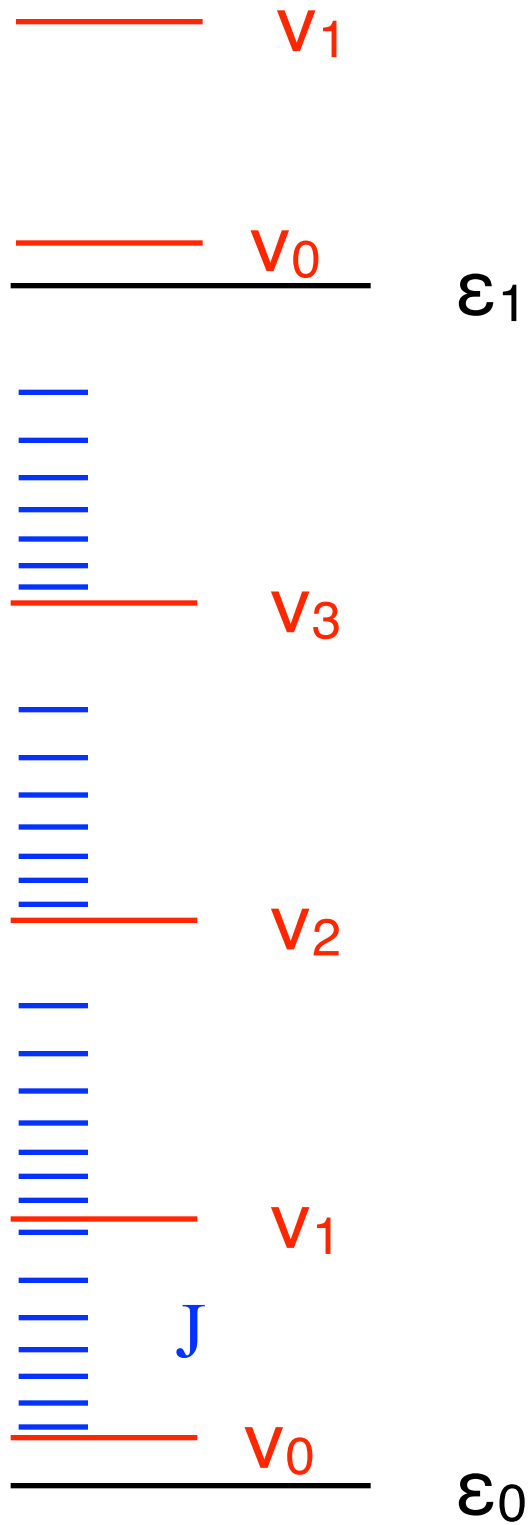
γ - half-width at half-maximum ($\text{cm}^{-1}/\text{atm}$)

δ - line shift ($\text{cm}^{-1}/\text{atm}$)

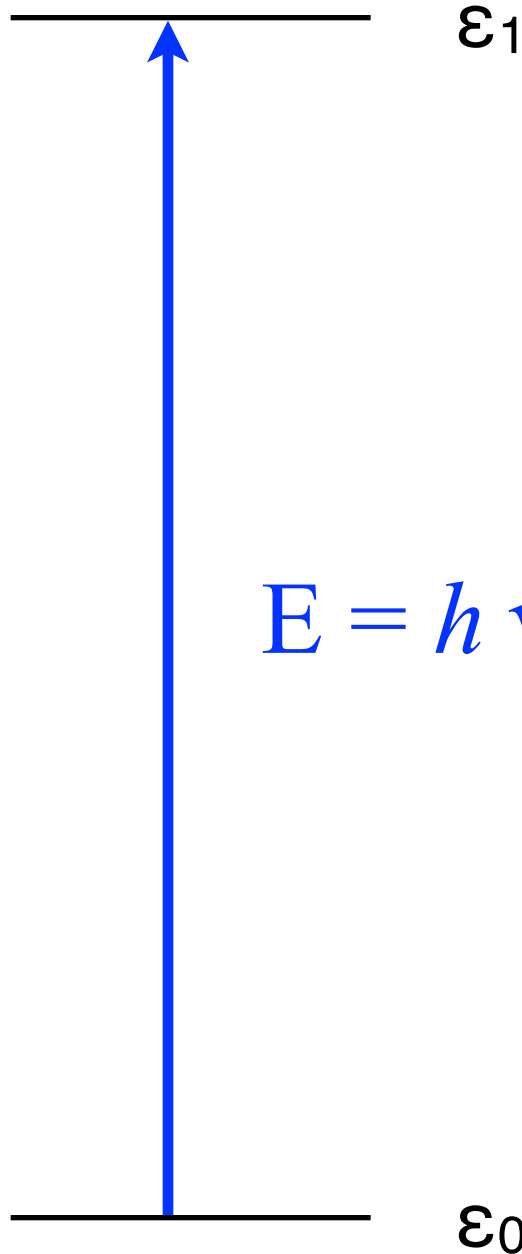
E'' - lower state energy (cm^{-1})

n - temperature dependence of γ

Energy Structure



Visible - UV region

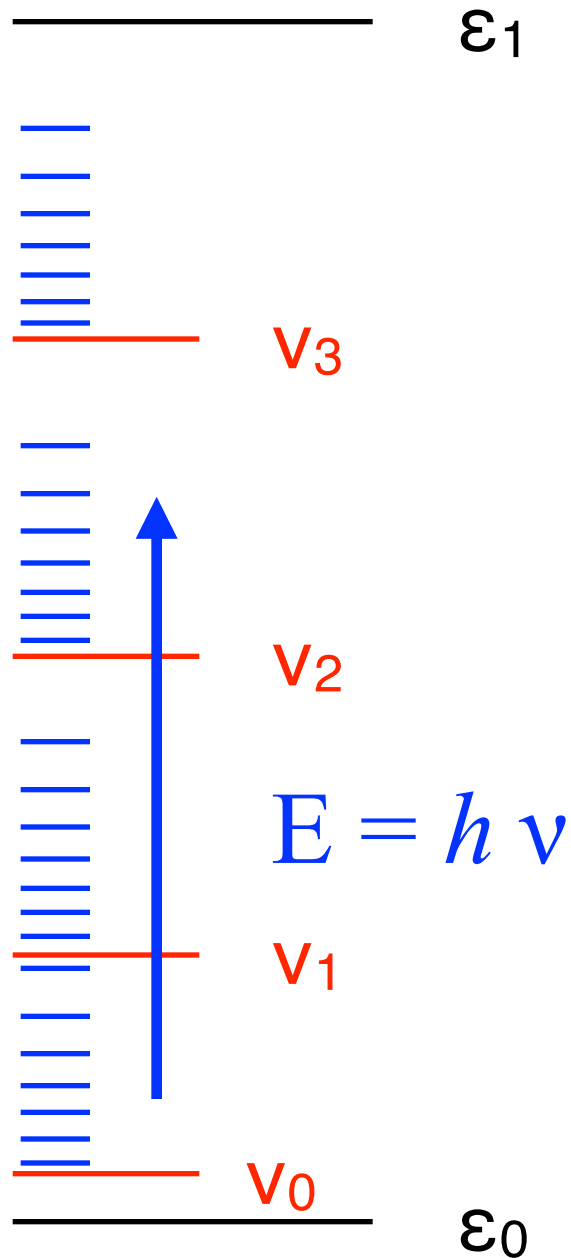


change in the electronic
state

high energy

$>14,000$ wavenumbers

Infrared region

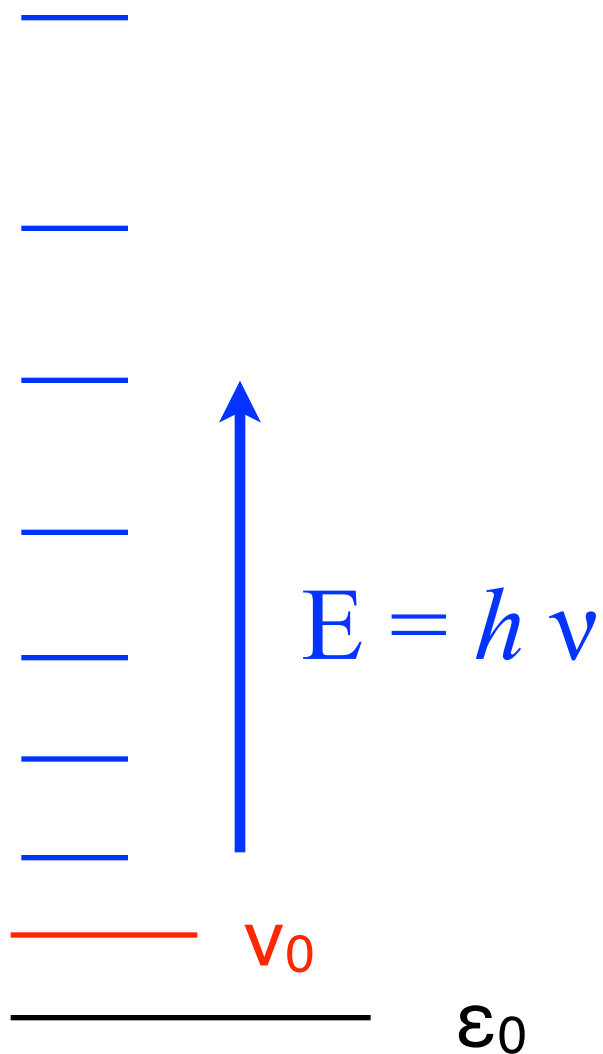


transitions are
between
rovibrational states

1000 to 14,000
wavenumbers

microwave region

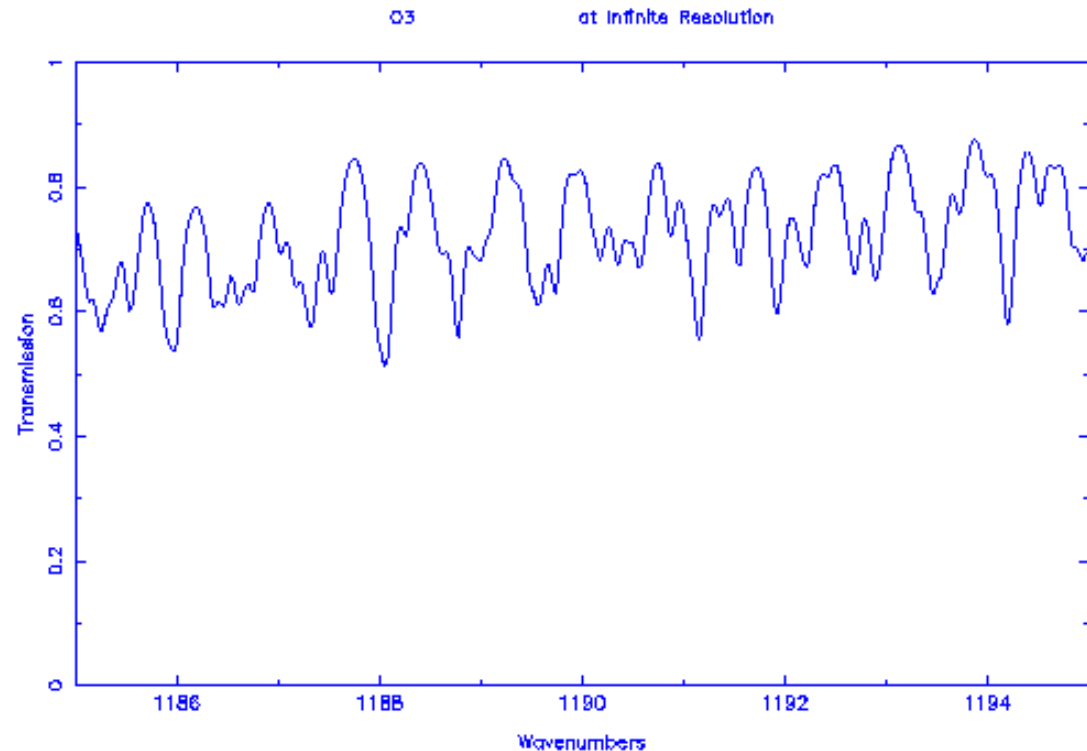
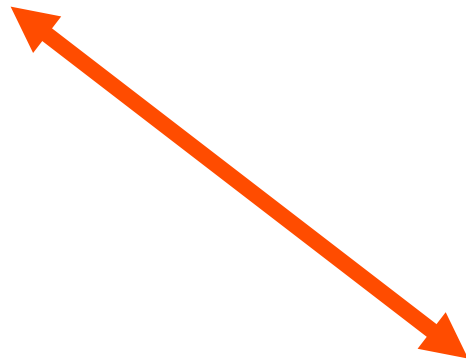
transitions are
between rotational
levels



0 to 1000
wavenumbers

Parameters \leftrightarrow Spectrum

Spectral Parameters



Spectral parameters come from:

- HITRAN Database
- GEISA Database

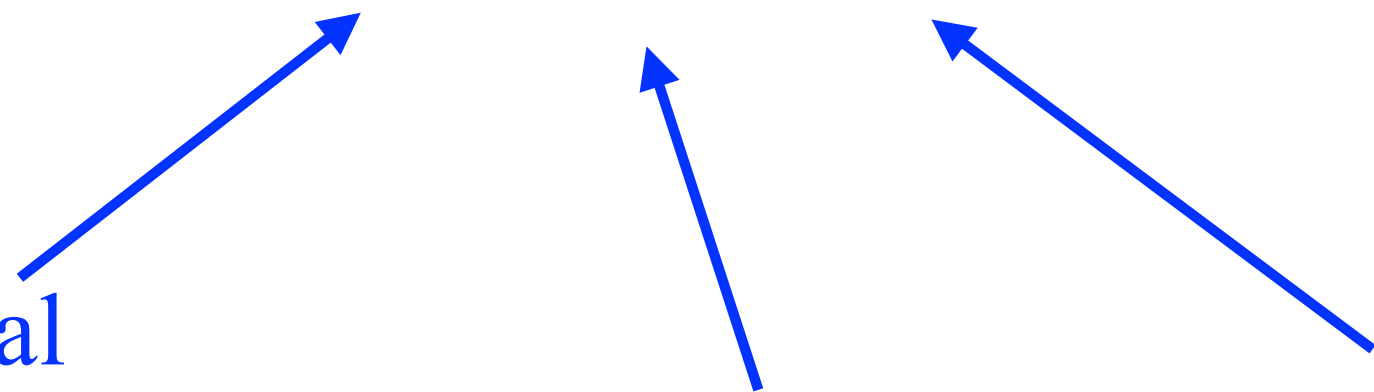
But data is incomplete and only approximate

Energy Structure

The Schrodinger Equation

$$H \Psi = E \Psi$$

Total
energy
operator
 $T+V$



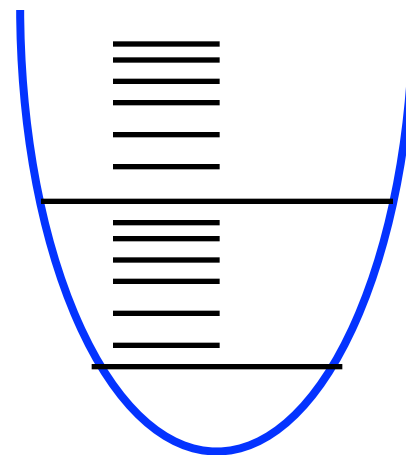
wavefunction

Energy

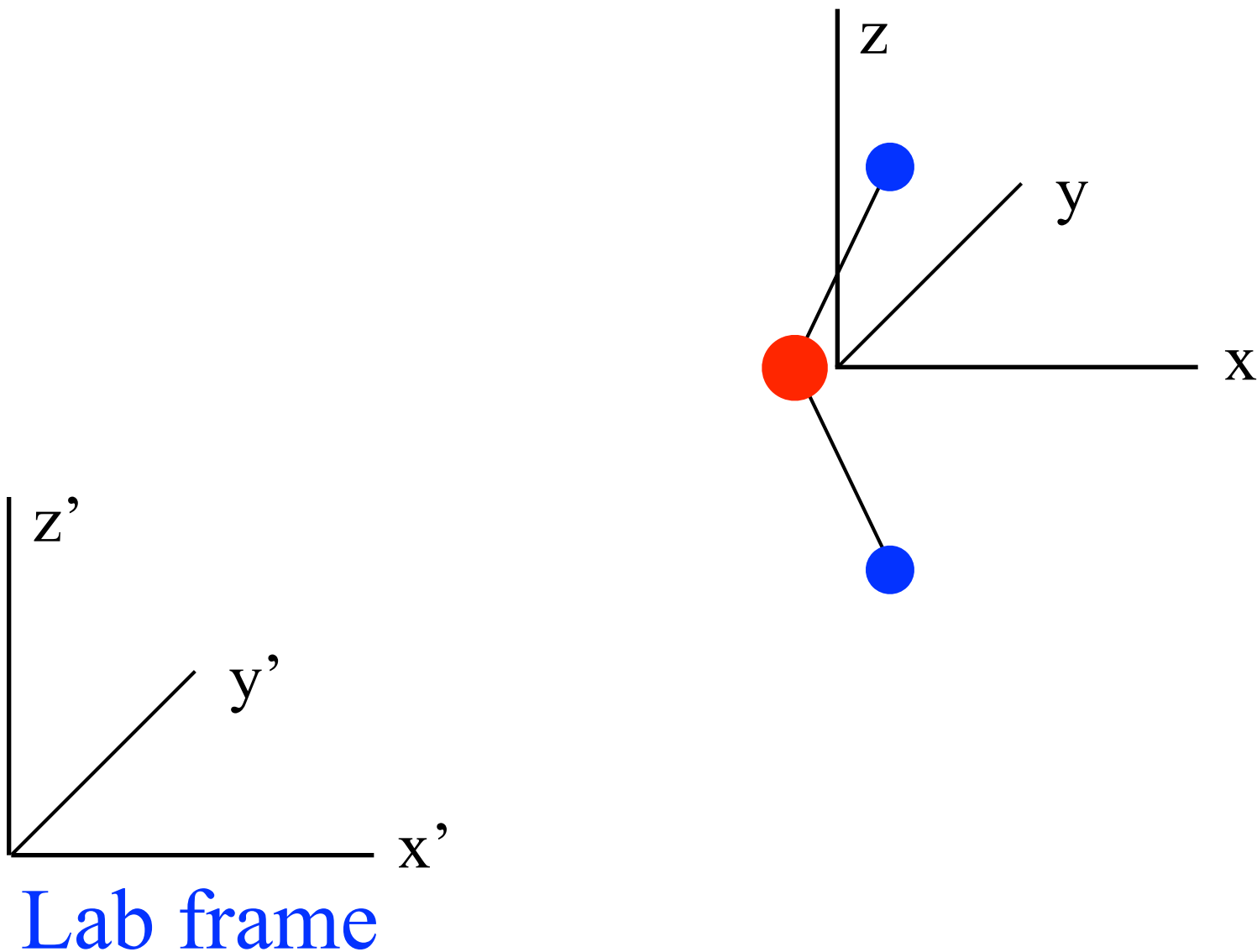
Rotational Energy Structure

$$H \Psi = E \Psi$$

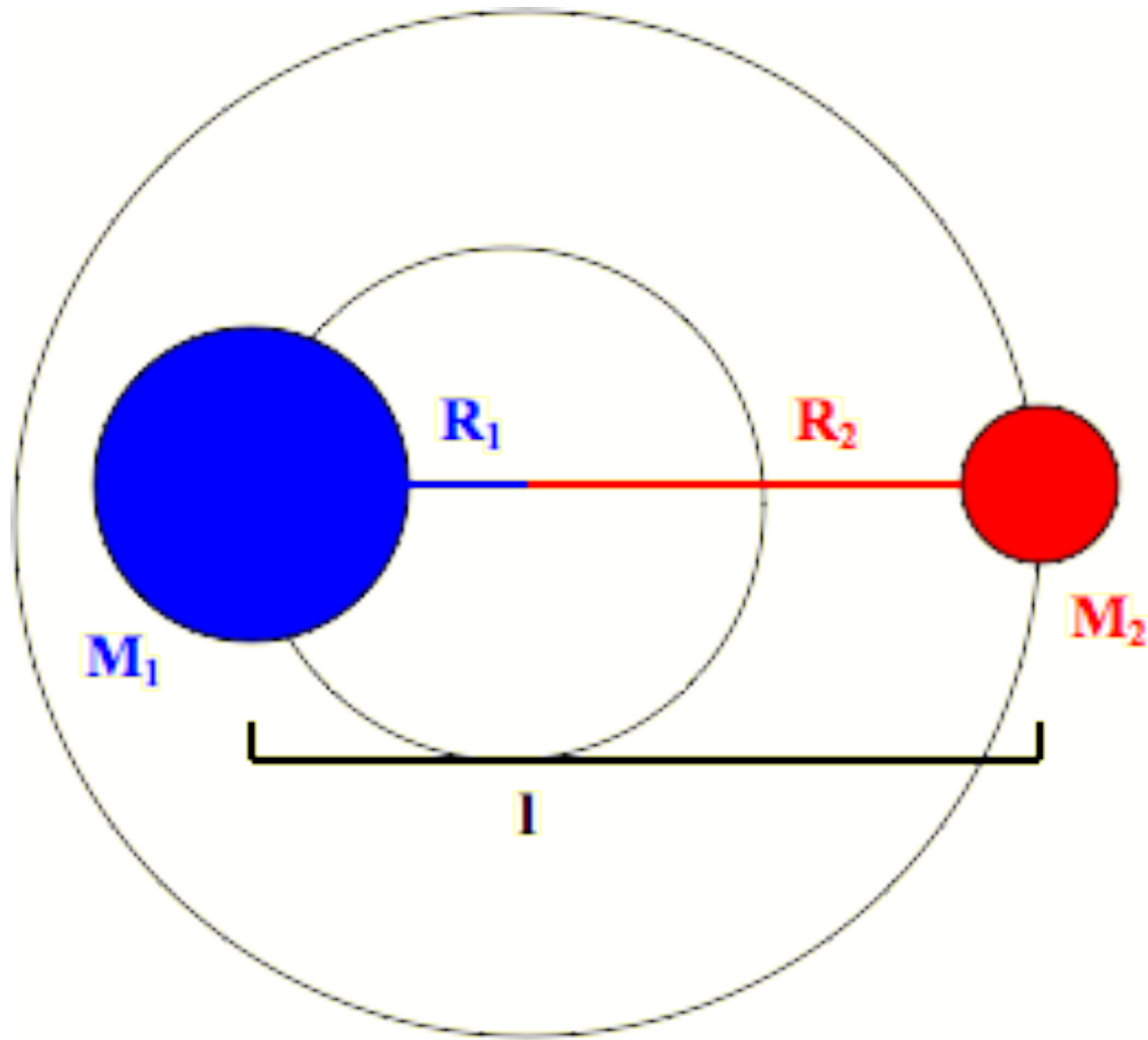
Here the kinetic energy is that of the molecule rotating and potential energy of the vibrational state.



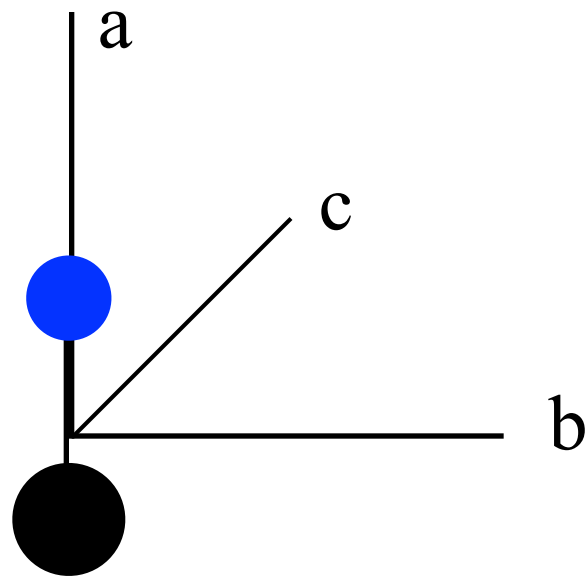
Lab vs. molecule fixed frames



where $I = \sum m_i r_i^2$, I is a tensor



Linear Molecule



$$I_A = 0 \text{ and } I_B = I_C$$

moment of inertia

The energy of a rotation molecule is given in terms of the moment of inertia

$$E(cm^{-1}) = \frac{1}{2hc} I_B \omega^2$$

$$E(cm^{-1}) = \frac{h}{8\pi^2 c I_B} J(J+1)$$

Linear Molecules

$$E = \frac{h}{8 \pi^2 c I_B} J(J+1)$$

we define the rotational constant B

$$B = \frac{h}{8 \pi^2 c I_B}$$

which gives

$$E = B J(J+1)$$

Determining the Rotational Constants

Consider the transitions $J \rightarrow J + 1$ and $J + 1 \rightarrow J + 2$.

The frequencies are given by

$$\nu_1 = E_{J+1} - E_J \text{ and } \nu_2 = E_{J+2} - E_{J+1} \text{ .}$$

The separation between the lines is

$$\Delta\nu = \nu_2 - \nu_1$$

$$E = B J (J + 1)$$

Determining the Rotational Constants

Consider the transitions from a level J to $J + 1$

$$\nu = E_{J+1} - E_J$$

$$\Delta\nu = B(J+1)(J+2) - BJ(J+1)$$

Determining the Rotational Constants

$$\Delta \nu = B(J^2 + 3J + 2 - J^2 - J)$$

$$\Delta \nu = 2B(J + 1)$$

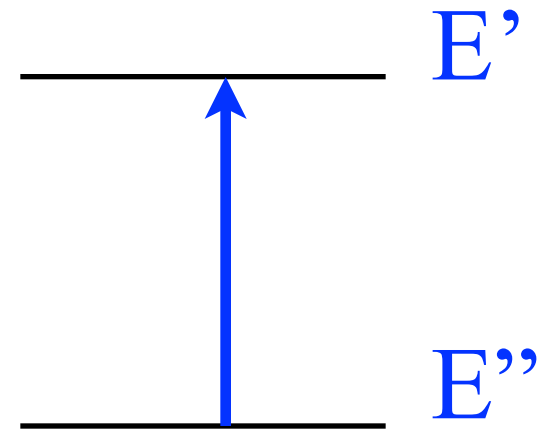
The spacing between adjacent lines is

Get the separation between lines ($\Delta\nu$)
from the spectra

$$\Delta\nu(J+1) - \Delta\nu(J) = 2B$$

Spectroscopic Convention

Consider a transition from some lower state to an upper state.



The frequency is given by $\nu = E' - E''$

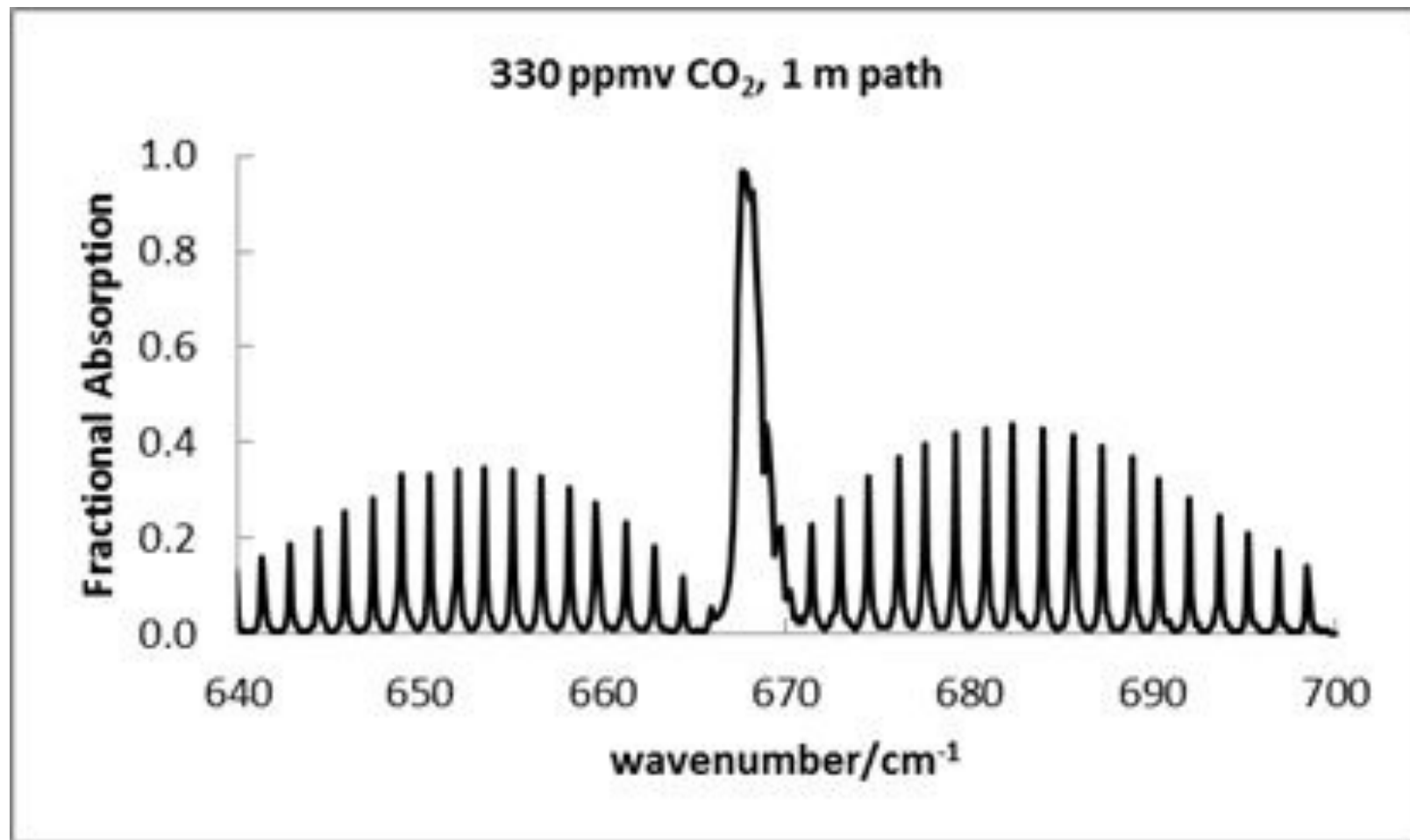
The spectroscopic convention is

‘ labels the upper state and

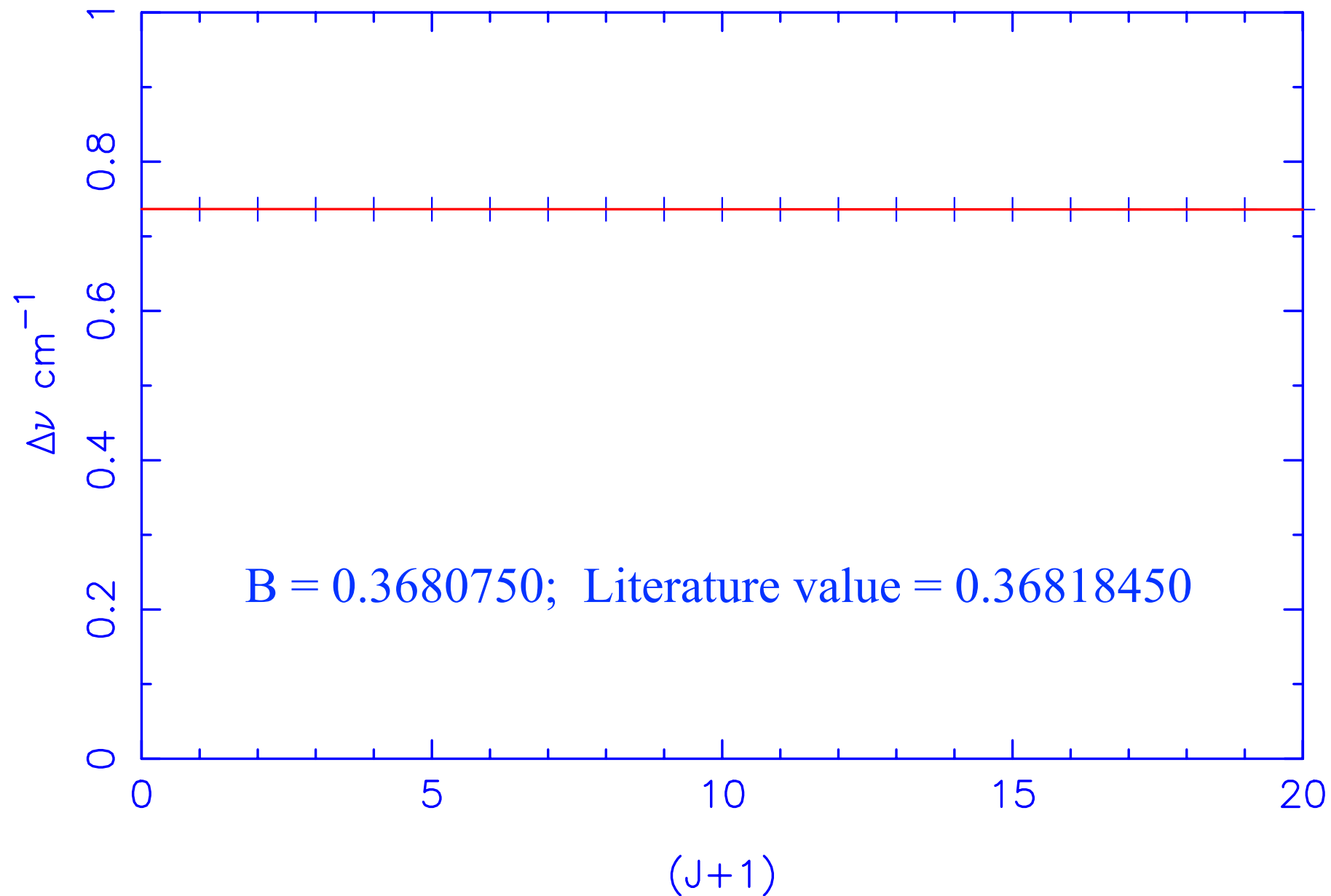
“ labels the lower state.



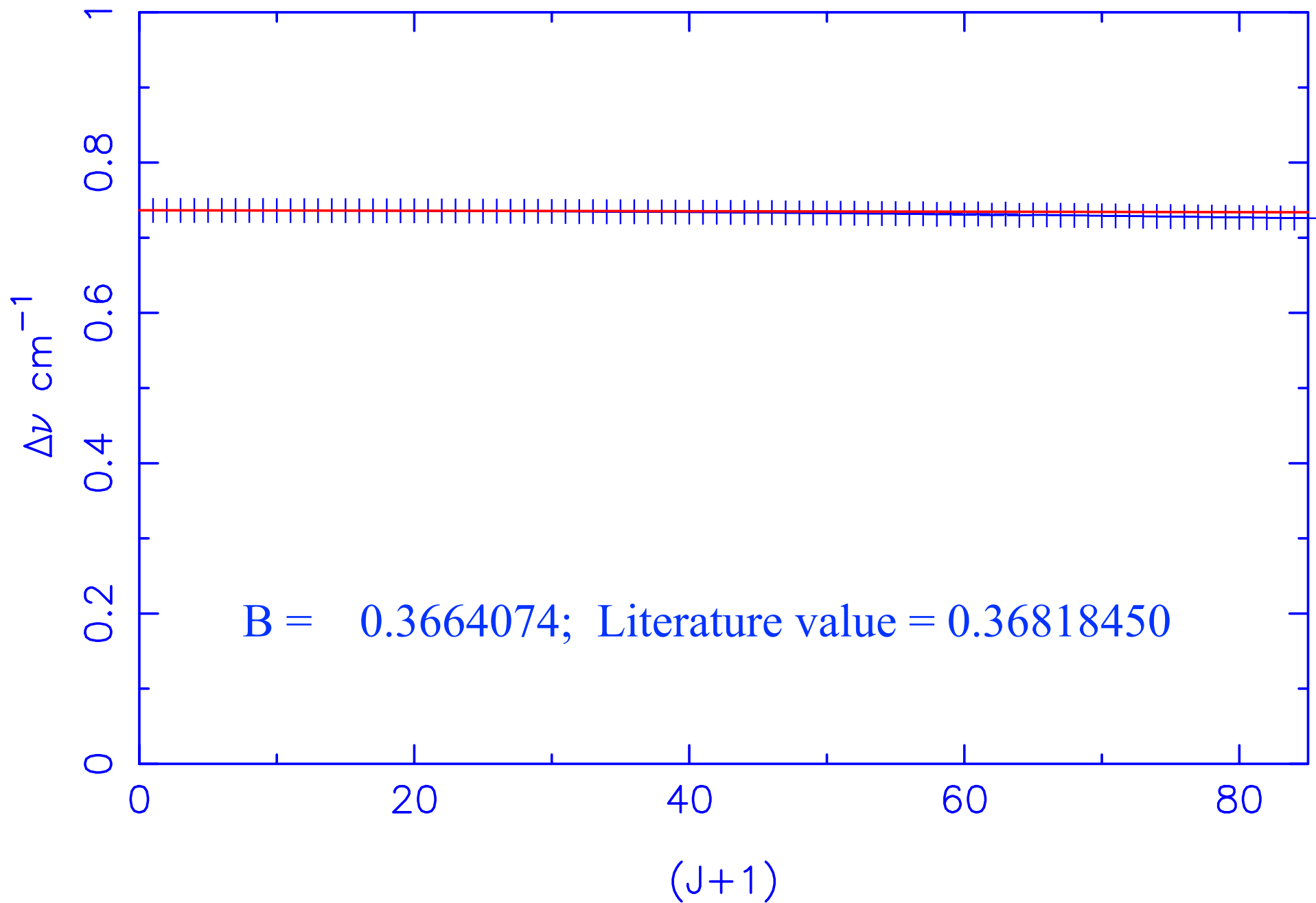
Get the separation between lines ($\Delta\nu$) from the spectra



$^{16}\text{O}^{12}\text{C}^{18}\text{O}$ rotation band

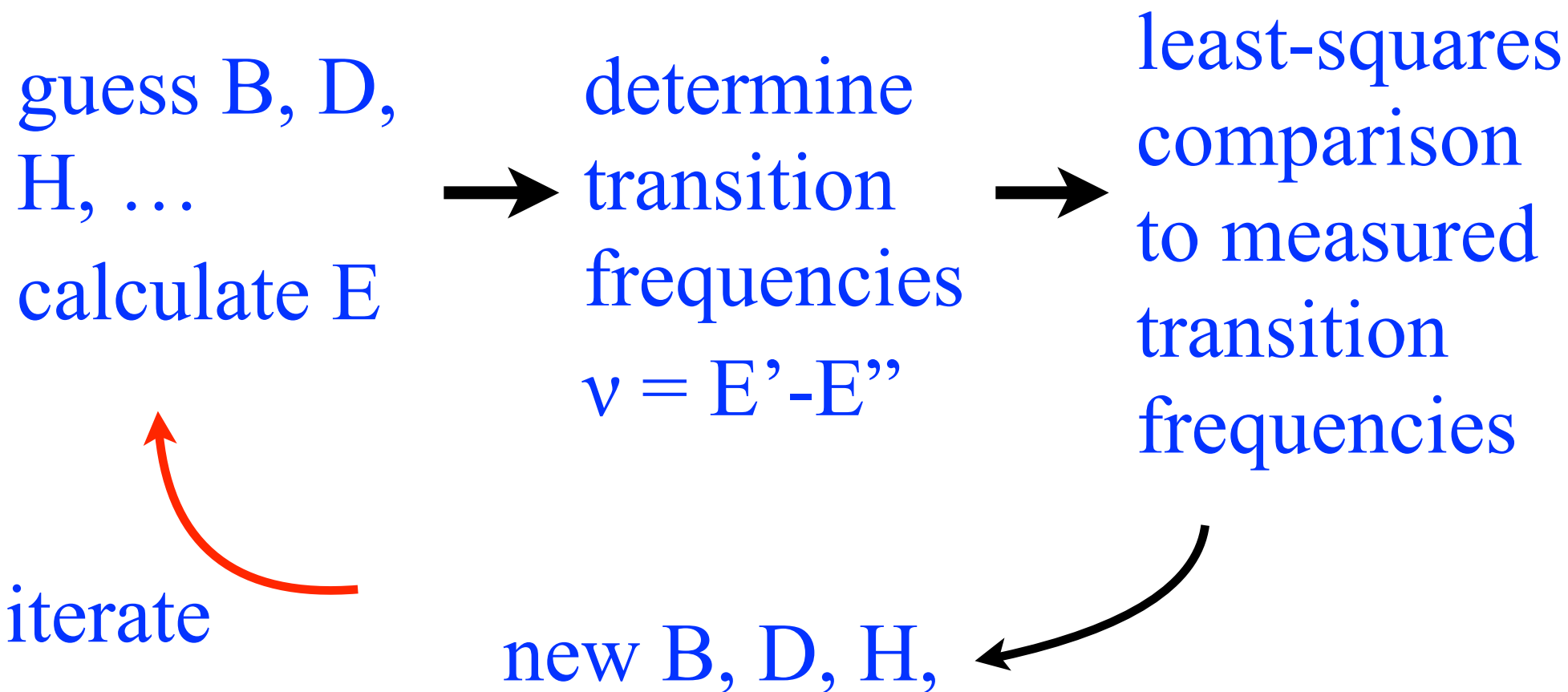


$^{16}\text{O}^{12}\text{C}^{18}\text{O}$ rotation band



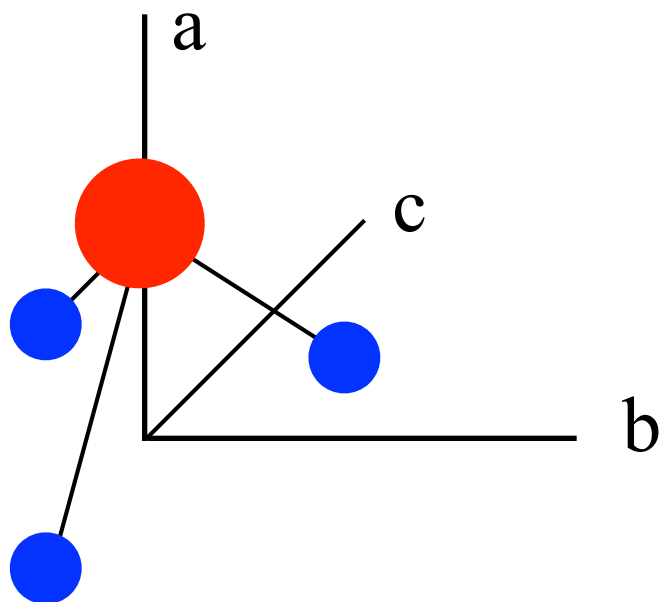
Rotational Energy of a Linear Molecule

$$E = B J(J+1) - D J^2(J+1)^2 + H J^3(J+1)^3 + \dots$$



Symmetric Top Molecules

Examples are NH_3 , CH_3Cl , PH_3



$$I_A \neq 0$$

$$I_B = I_C \neq 0$$

Energy - Symmetric Top

$$A = \frac{h}{8\pi^2 c I_A} \qquad B = \frac{h}{8\pi^2 c I_B}$$

$$E(J, K) = B J(J+1) + (A - B) K^2$$

J is the total angular momentum quantum number $J = 0, 1, 2, \dots$

K is the component of angular momentum about the figure axis:

$K = J, J-1, J-2, \dots -J$

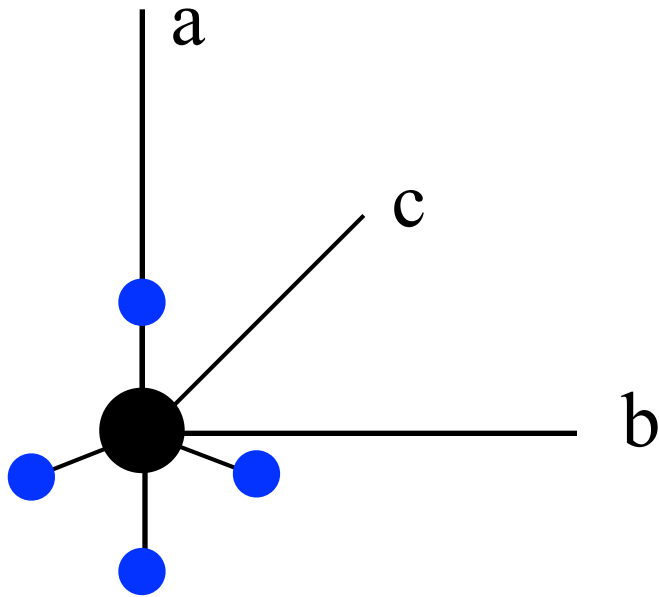
Energy - Symmetric Top

$$E(J, K) = B J(J+1) + (A - B) K^2$$

when Coriolis and other resonances are accounted for the energy expression is much more complicated. See Poynter RL, Margolis JS. The ground state far infrared spectrum of NH₃. Mol Phys 1983;48:401-18.

Spherical Top Molecules

Examples are CH_4 , CCl_4 , SF_6



$$I_A = I_B = I_C$$

Energy - Symmetric Top

$$B = \frac{h}{8 \pi^2 c I_B}$$

$$E(J) = B J(J+1)$$

J is the total angular momentum quantum number $J = 0, 1, 2, \dots$

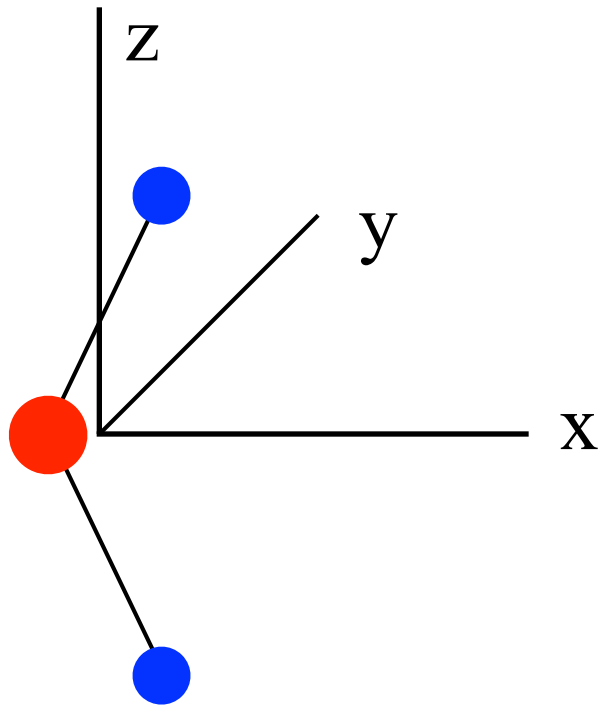
Note inclusion of resonances gives a much more complicated Hamiltonian.

see

Jean-Paul Champion, Gérard Pierre, Hubert Berger, Jean Cadot, Vibration-rotation energies of harmonic and combination levels in tetrahedral XY_4 molecules: Analysis of the $2\nu_2$ and $\nu_2 + \nu_4$ bands of $^{12}\text{CH}_4$, Journal of Molecular Spectroscopy 79, 281–294 (1980).

Asymmetric top molecules

Examples are H_2O , HNO_3 , H_2CO , O_3 , ...



$$I_A \neq I_B \neq I_C \neq 0$$

Energy - Asymmetric Top

$$A = \frac{h}{8 \pi^2 c I_A} \quad B = \frac{h}{8 \pi^2 c I_B} \quad C = \frac{h}{8 \pi^2 c I_C}$$

$$E(J, K_a, K_c) = \text{complicated}$$

Flaud J-M, Camy-Peyret C. Vibration-Rotation intensities in H₂O- Type molecules application to the 2ν₂, ν₁, ν₃ Bands of H₂¹⁶O, J. Mol. Spectrosc. 1975; 55: 278-310