

# Skoog – Chapter 1

## Introduction

- Basics of Instrumental Analysis
  - Properties Employed in Instrumental Methods
  - Numerical Criteria
  - Figures of Merit

# Skip the following chapters

- Chapter 2 – Electrical Components and Circuits
- Chapter 3 – Operational Amplifiers in Chemical Instrumentation
- Chapter 4 – Digital Electronics and Microcomputers

# Skoog – Chapter 5

## Signals and Noise

- Signal to Noise Ratio

All instrumental measurements involve a signal

Unfortunately they always have noise present

Sometimes the noise is large

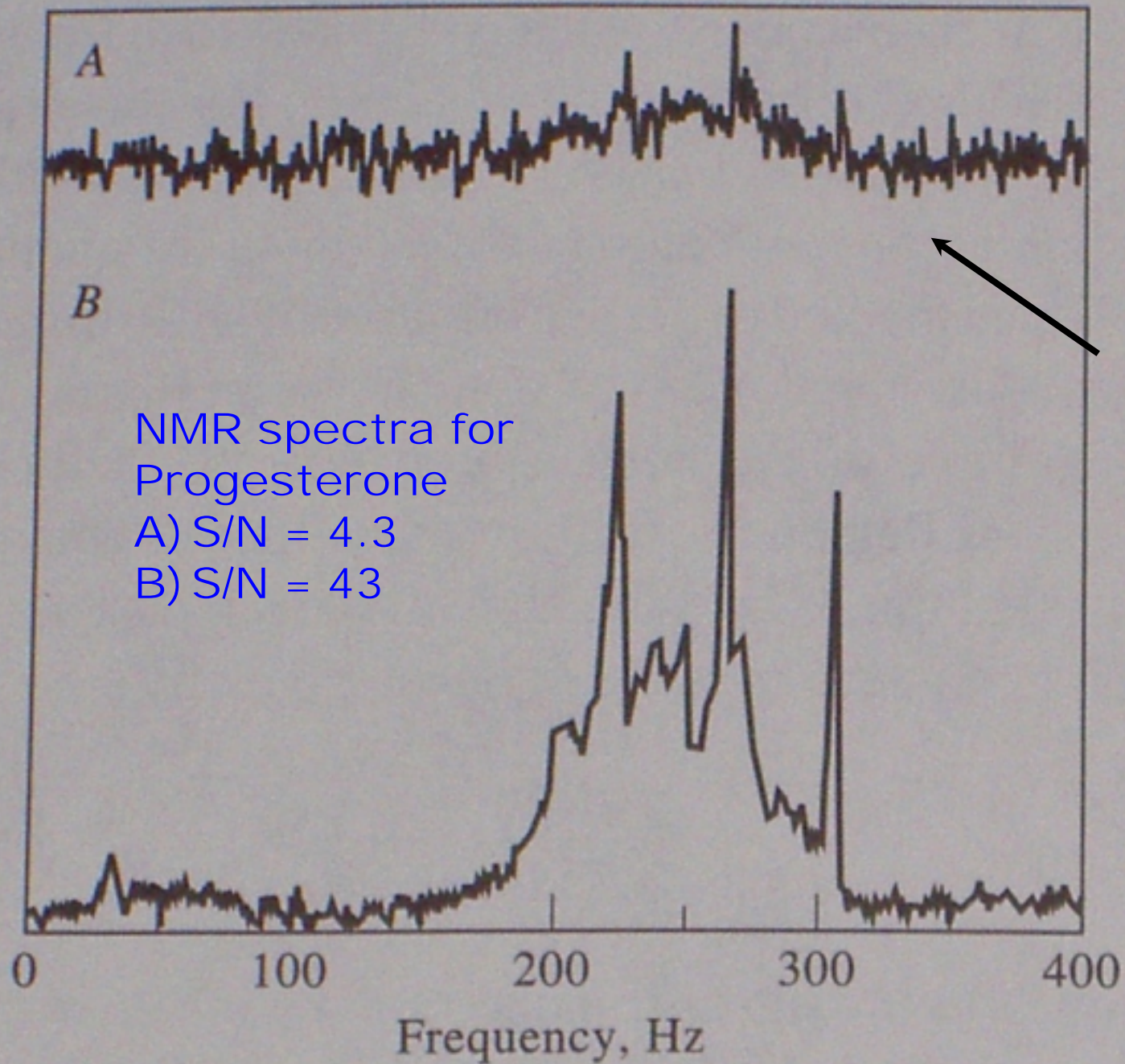
Sometimes it is so small you can't see it

# Signal to Noise Ratio (S/N)

- Parameter describing quality of data
- Often referred to as “figure of merit”

$$\frac{S}{N} = \frac{\text{mean of signal}}{\text{standard deviation}} = \frac{\bar{x}}{s} = \frac{1}{\text{RSD}}$$

RSD = relative standard deviation



Very little confidence in ability to determine peaks at lower S/N

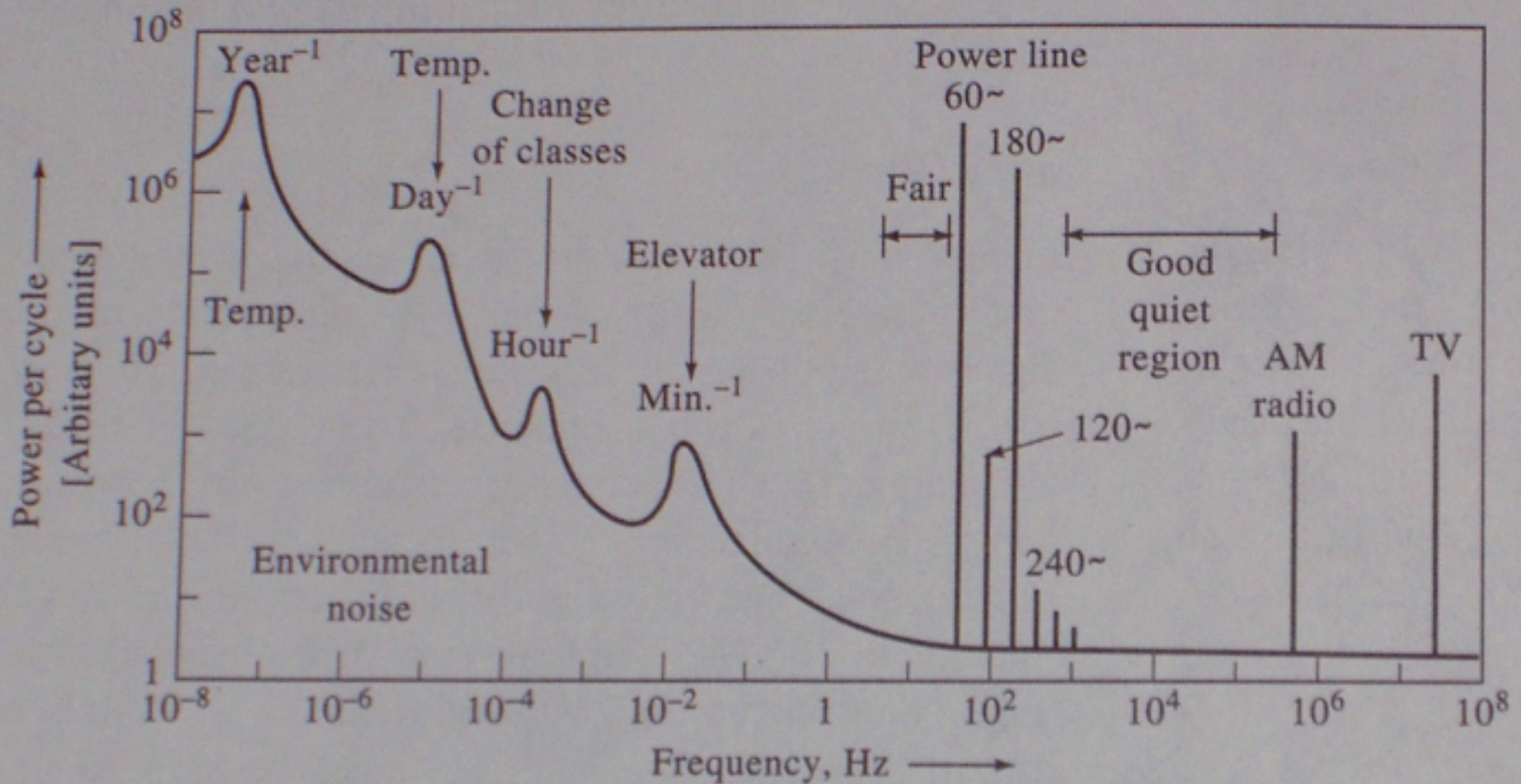
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Detection Limit occurs at S/N ~ 2 or 3

# Sources of Noise

- Chemical noise – temp, pressure, humidity, etc. fluctuations = uncontrolled variables
- Instrumental noise – noise from instrumental components
  - Thermal noise (Johnson noise) – thermal motion of electrons in load resistor
  - Shot noise – movement of electrons across a junction
  - Flicker noise – any noise that is inversely proportional to signal  $1/f$
  - Environmental noise – many noise sources

# Environmental noise sources (note frequency dependence)

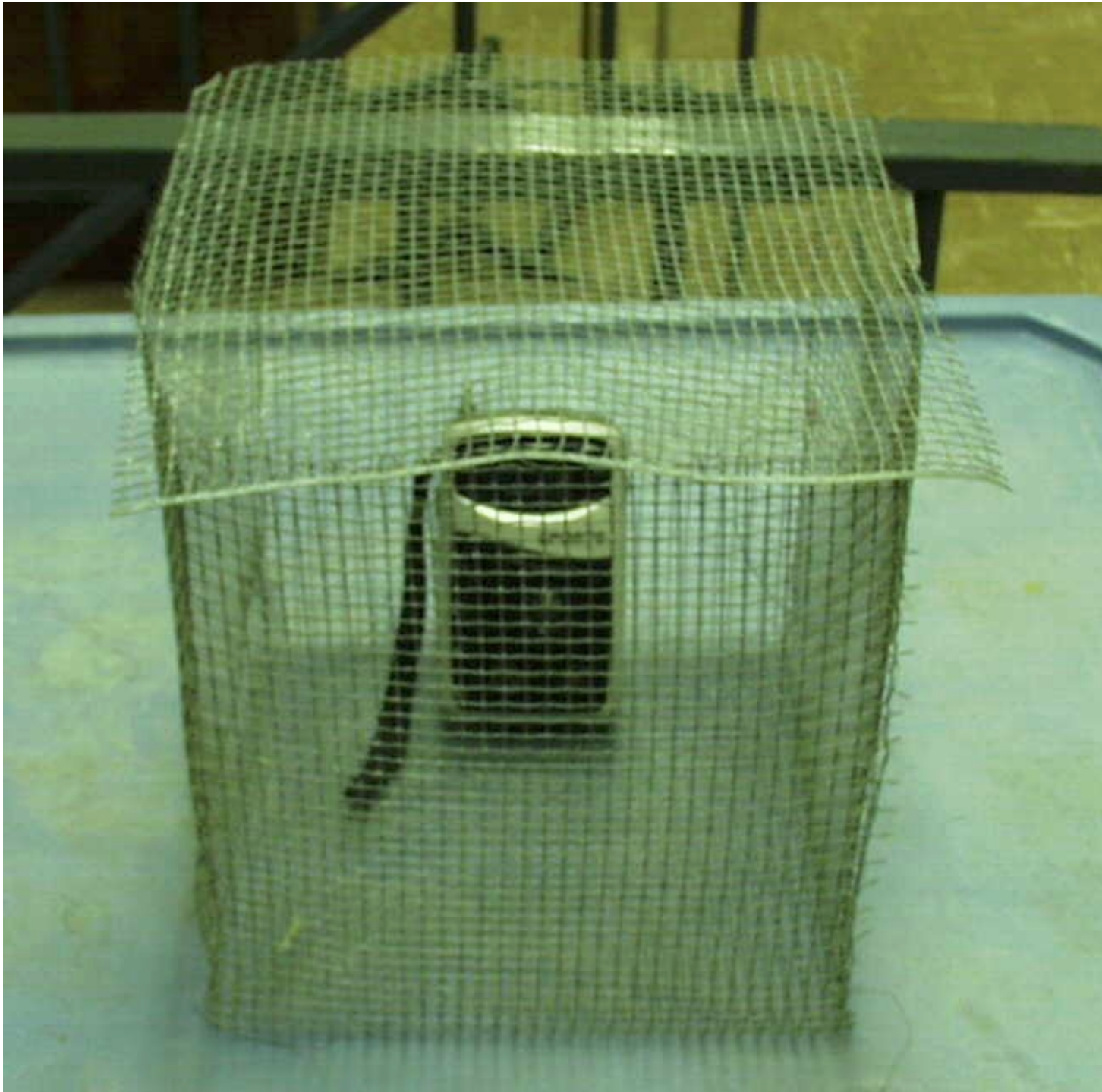


# Improving S/N

## hardware & software

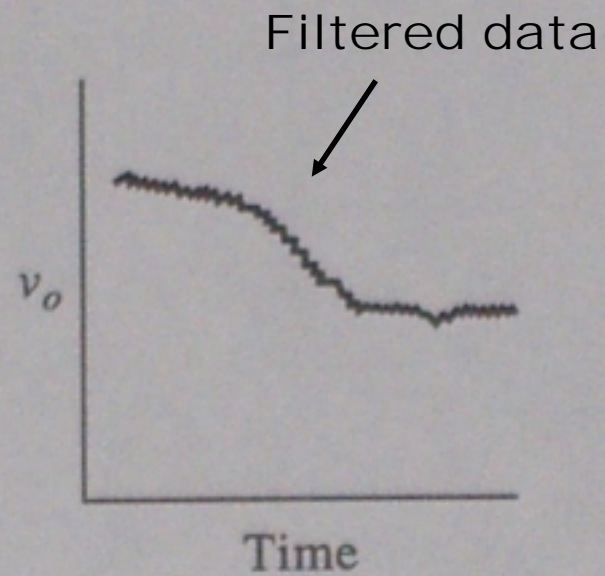
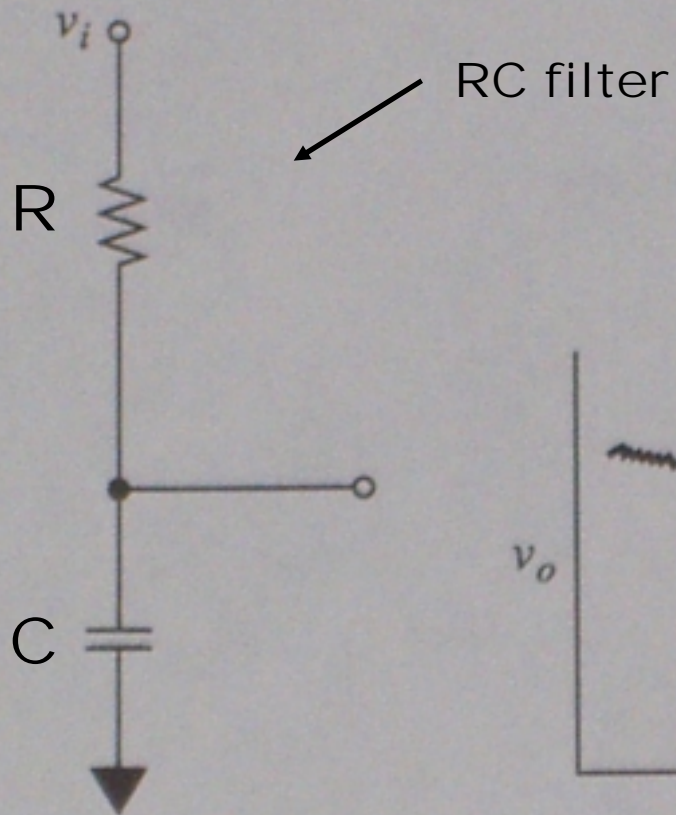
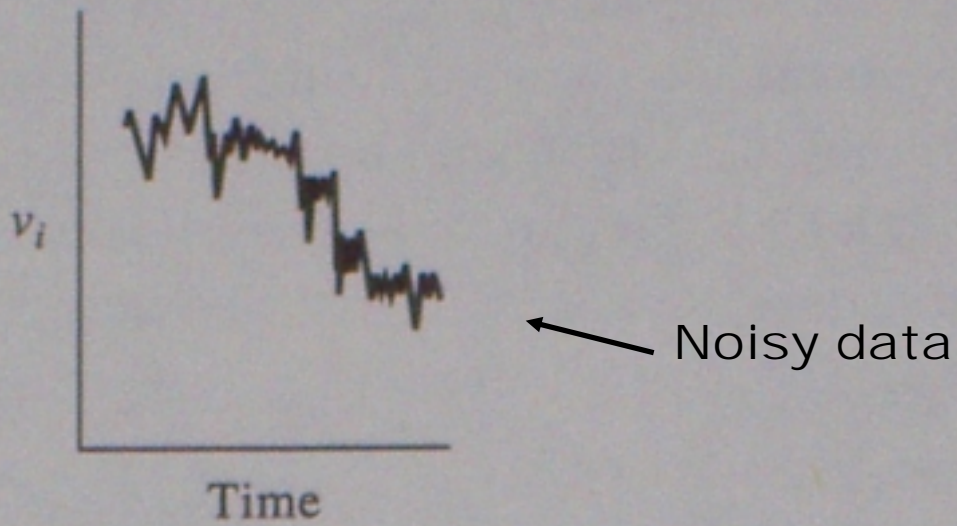
- Hardware
  - Grounding & shielding – Faraday cage
  - Analog filtering – RC filtering
  - Modulation – convert DC signal to high frequency AC then demodulate
  - Signal chopping – rotating wheel to differentiate e.g. IR source from heat
  - Lock-in amplifiers



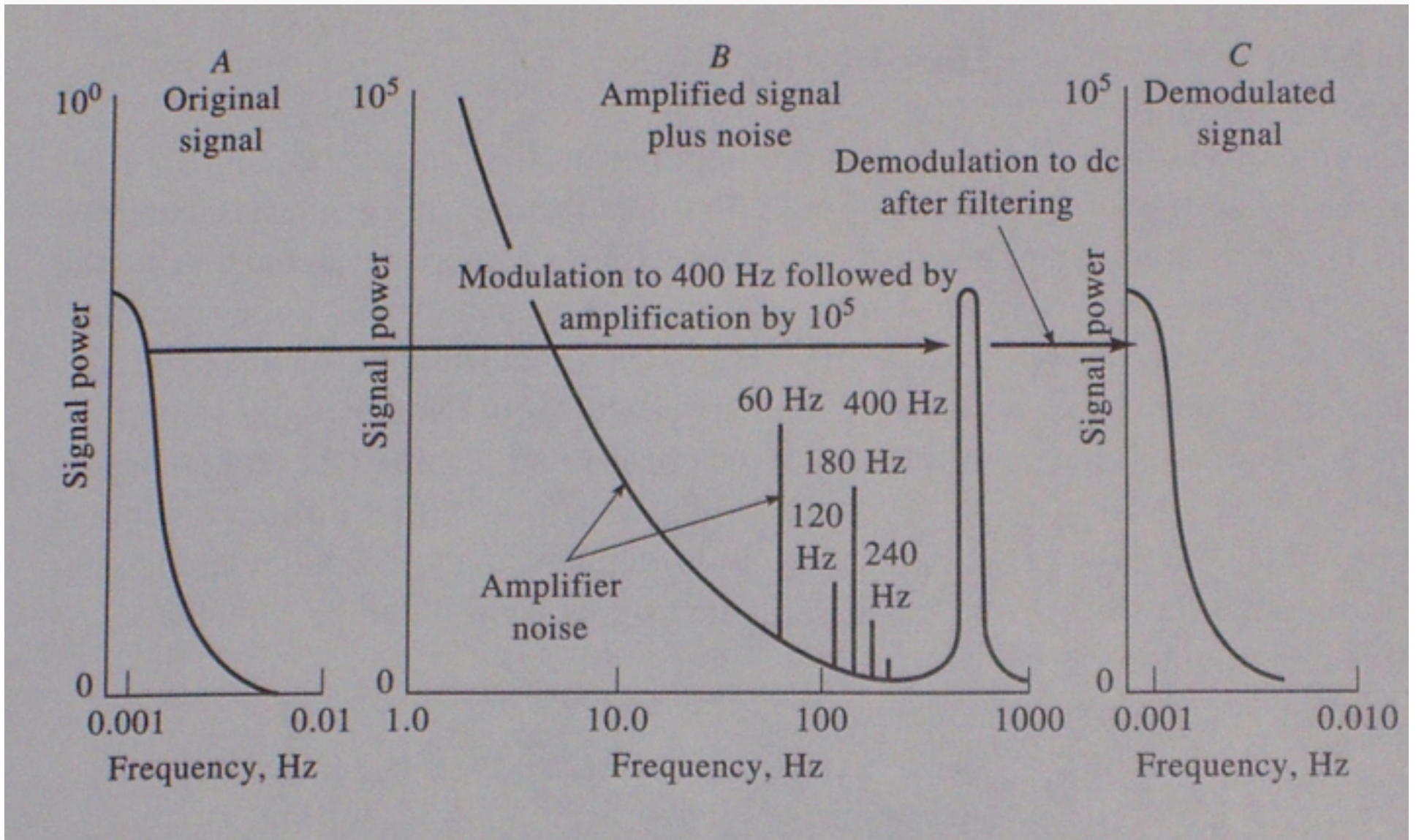


Primitive  
Faraday  
Cage

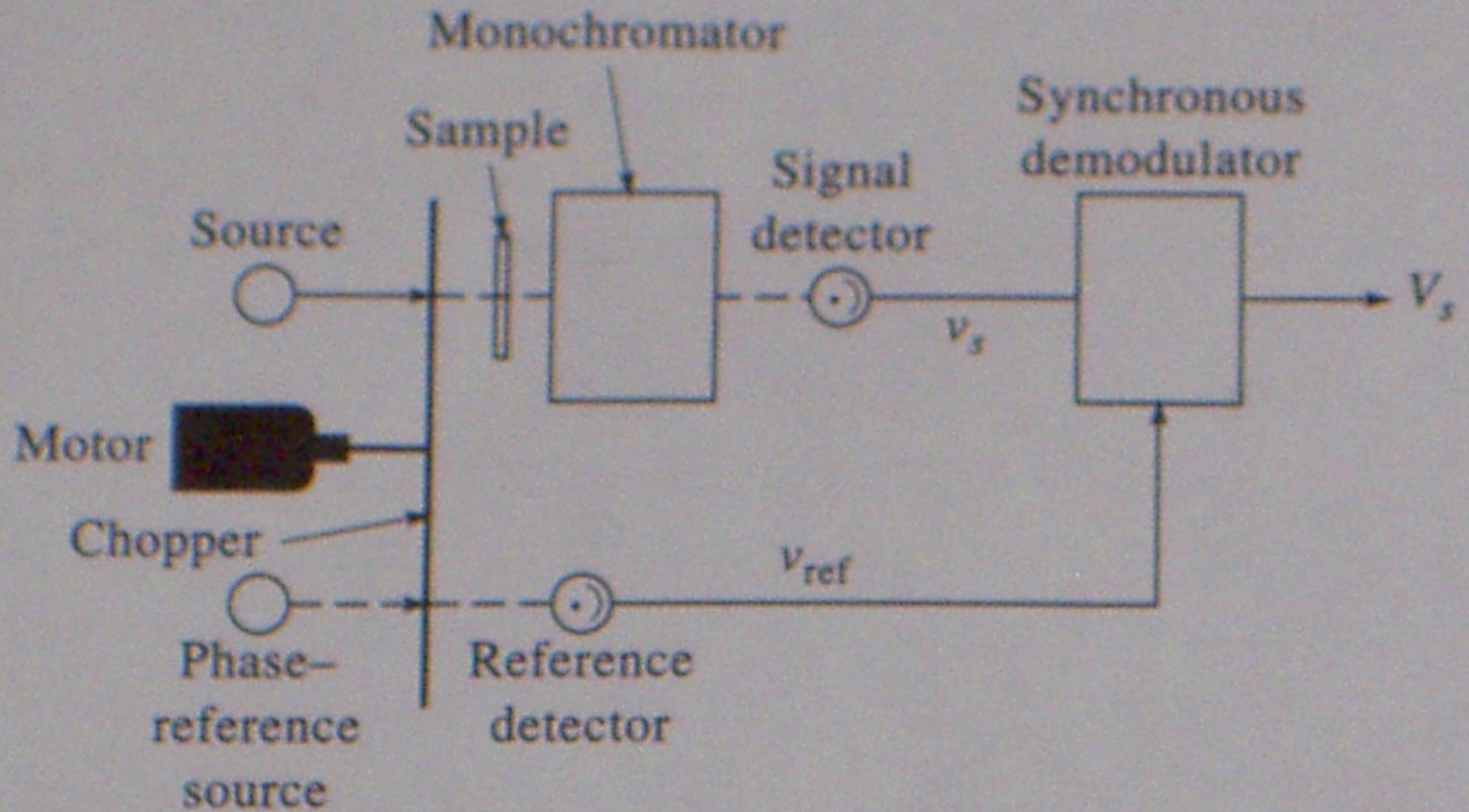
# Analog Filtering or RC Filtering

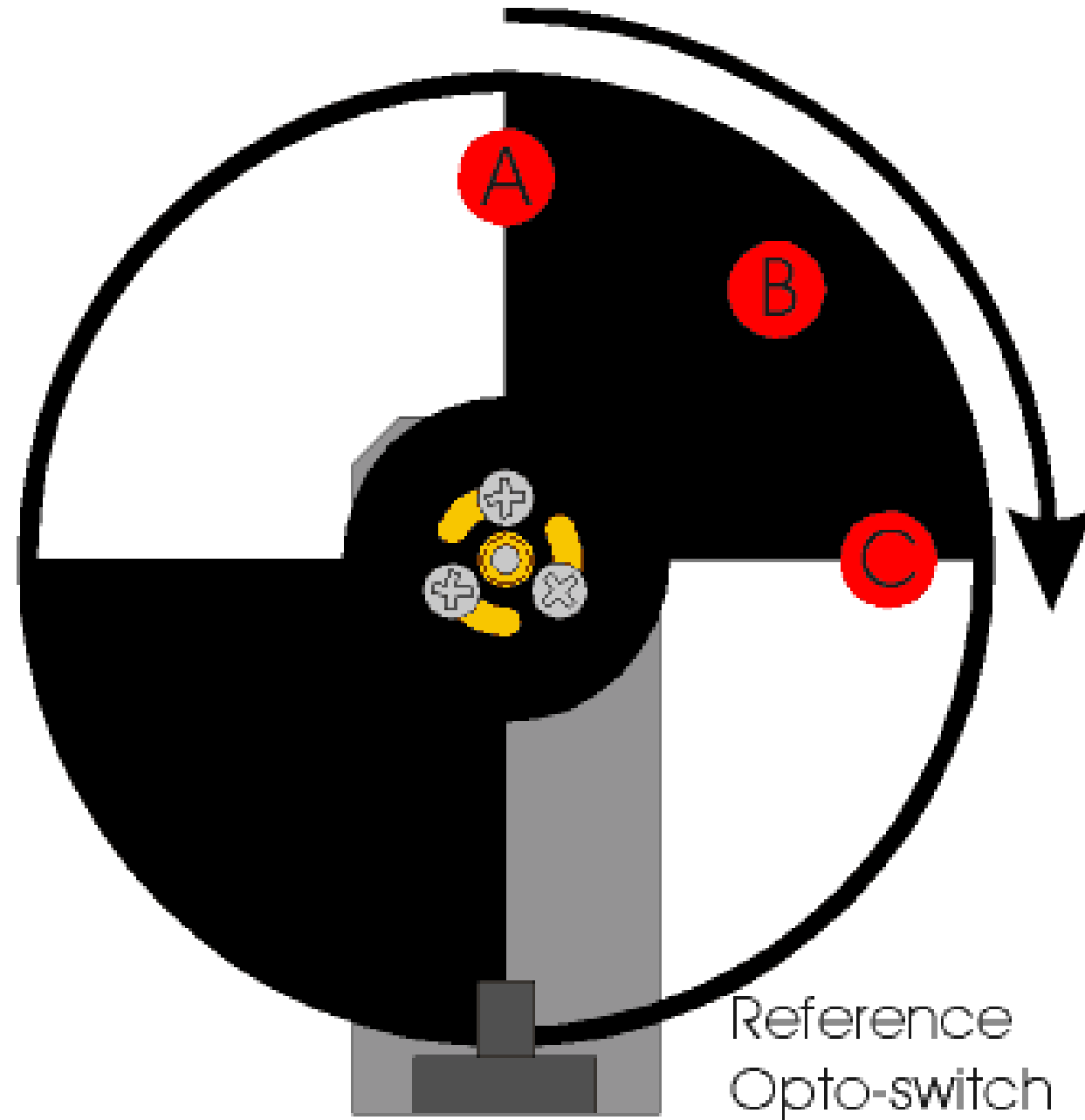


# Modulation



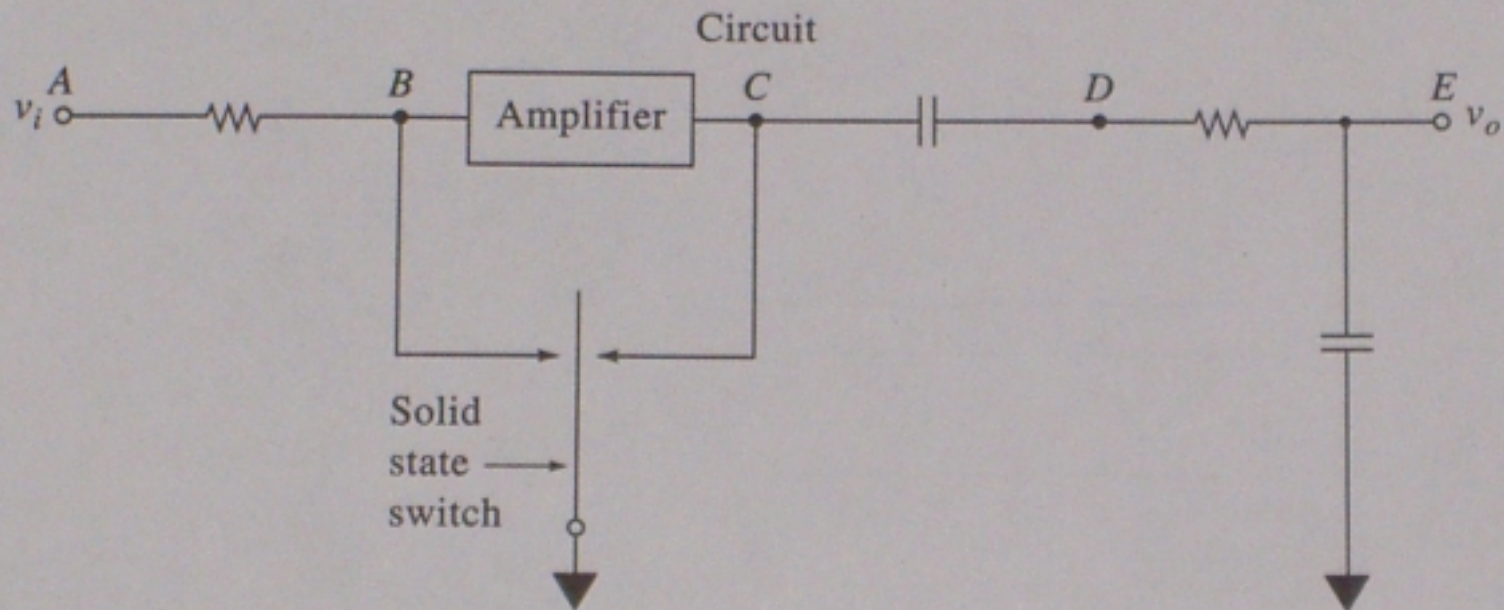
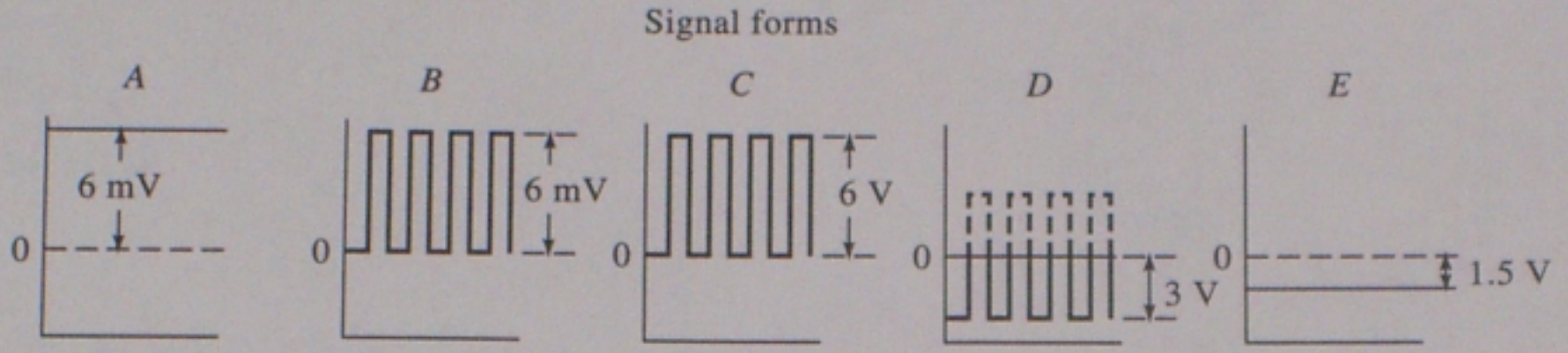
# Signal chopping in an IR spectrophotometer





Rotating  
Chopper

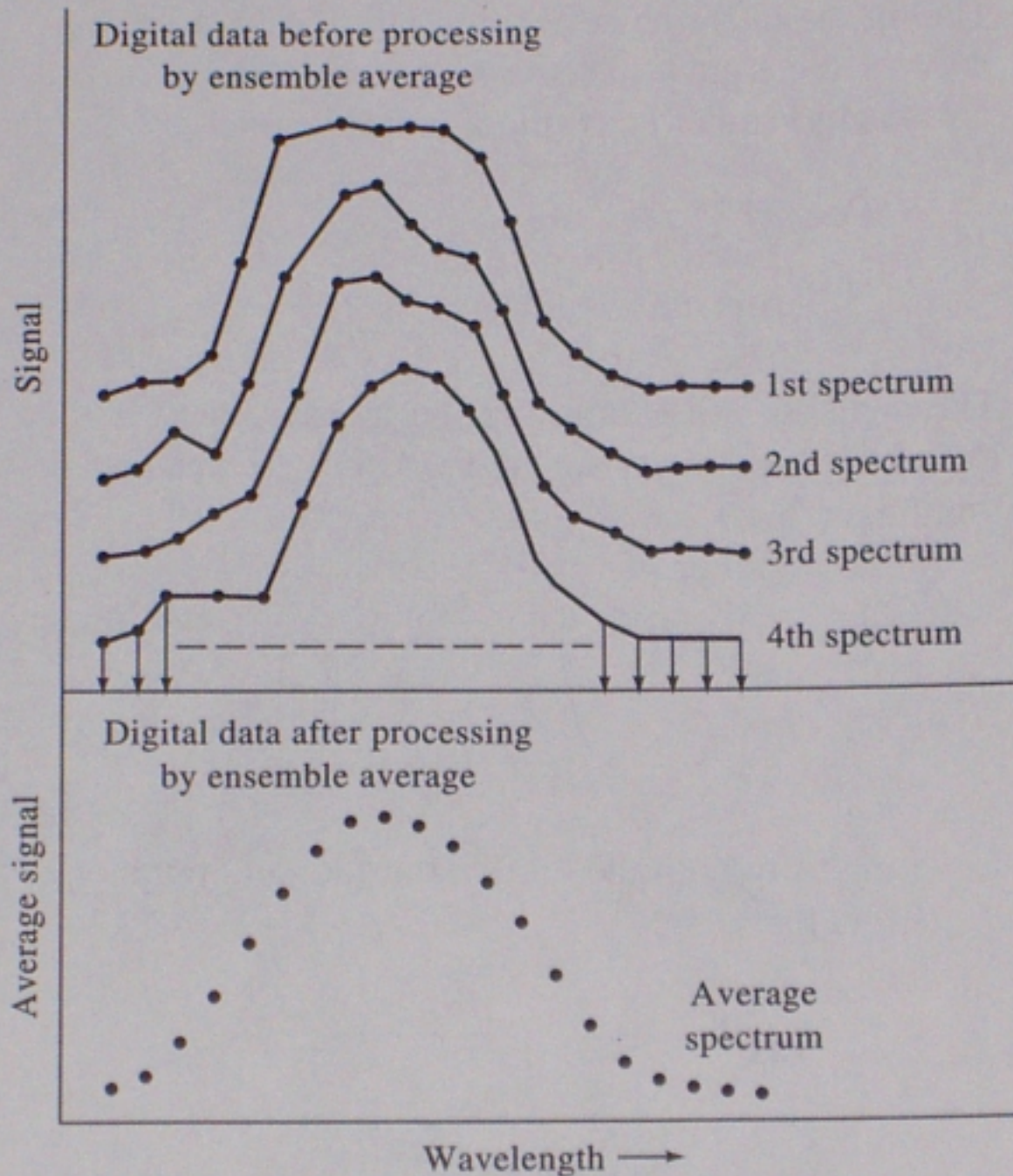
# Chopper amplifier



# Improving S/N

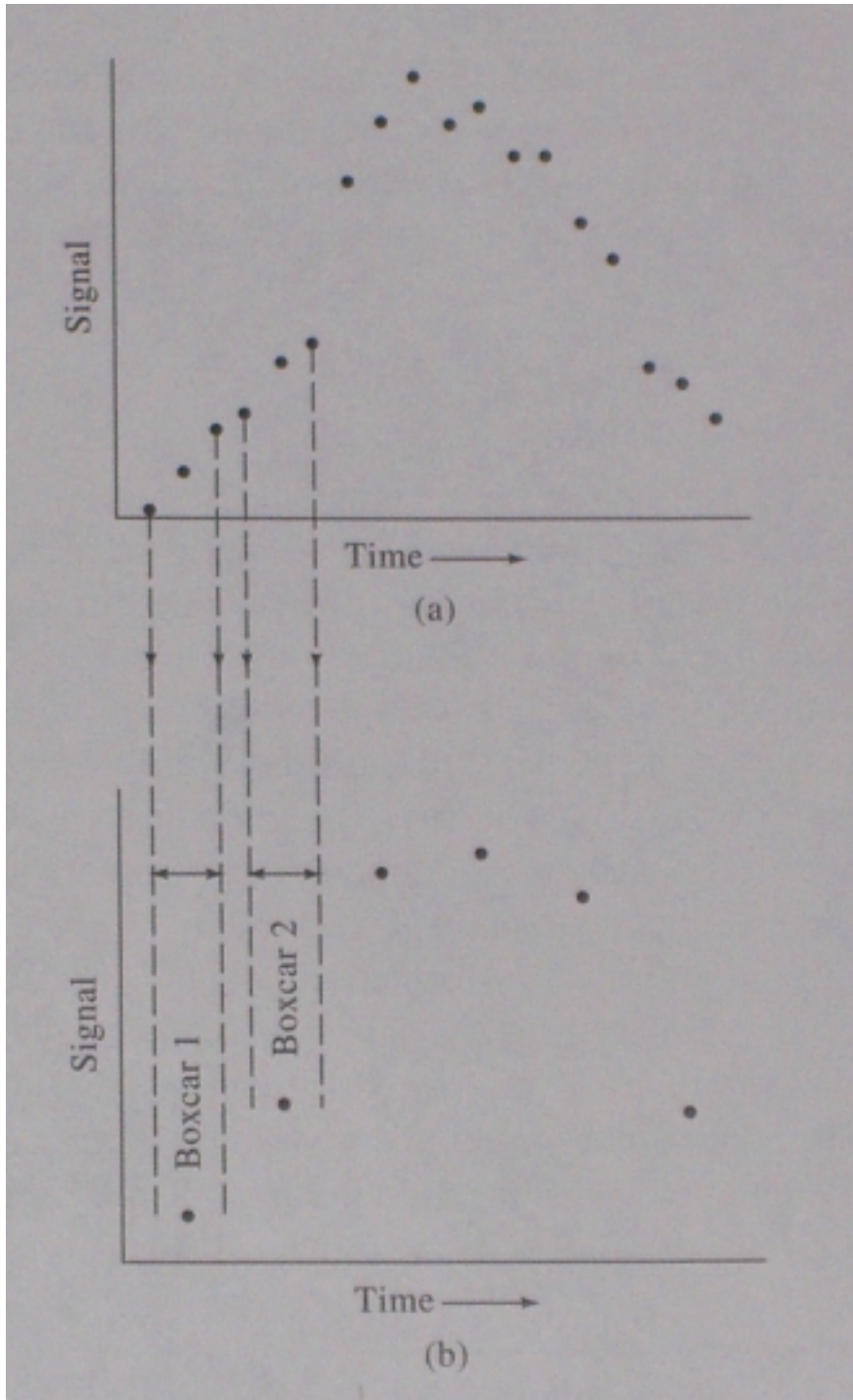
## hardware & software

- Software
  - Ensemble averaging – adding spectra
  - Boxcar averaging –
  - Digital filtering – moving window, sliding average
  - Correlation methods



Ensemble averaging  
i.e. adding  
or averaging  
signal





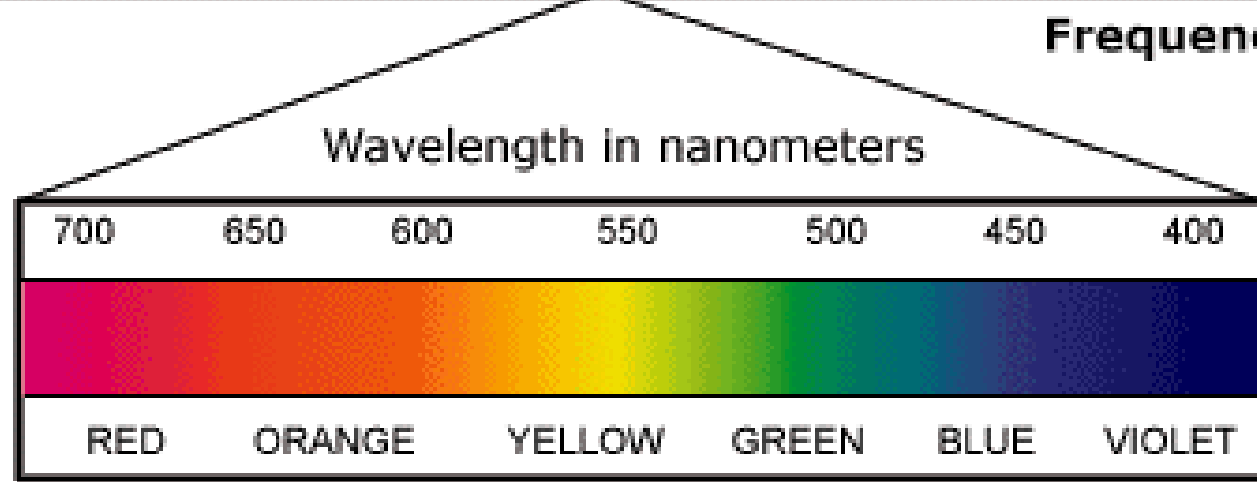
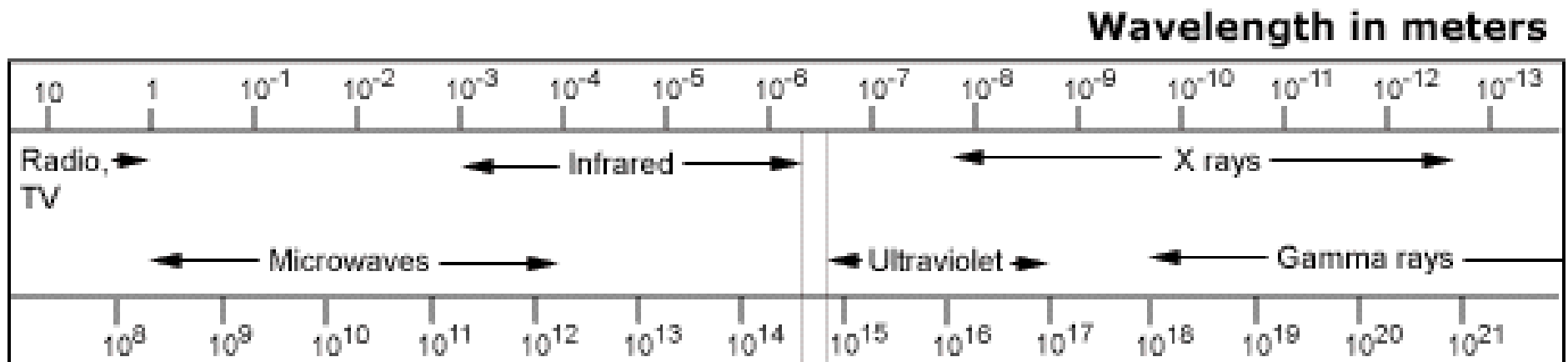
Boxcar averaging

# Skoog – Chapter 6

## Intro to Spectrometric Methods

- General Properties of Electromagnetic Radiation (EM)
- Wave Properties of EM
- Quantum-Mechanical Properties of EM
- Quantitative Aspects of Spectrochemical Measurements

Spin States	Molecular Rotations	Molecular Vibrations	Outer Shell Electrons	Inner Shell Electrons	Nuclear Transitions
NMR EPR	Microwave Absorption Spectroscopy	Infrared Absorption Spectroscopy	UV-vis Absorption, Fluorescence	X-Ray Absorption, Fluorescence	Gamma Ray Spectroscopy

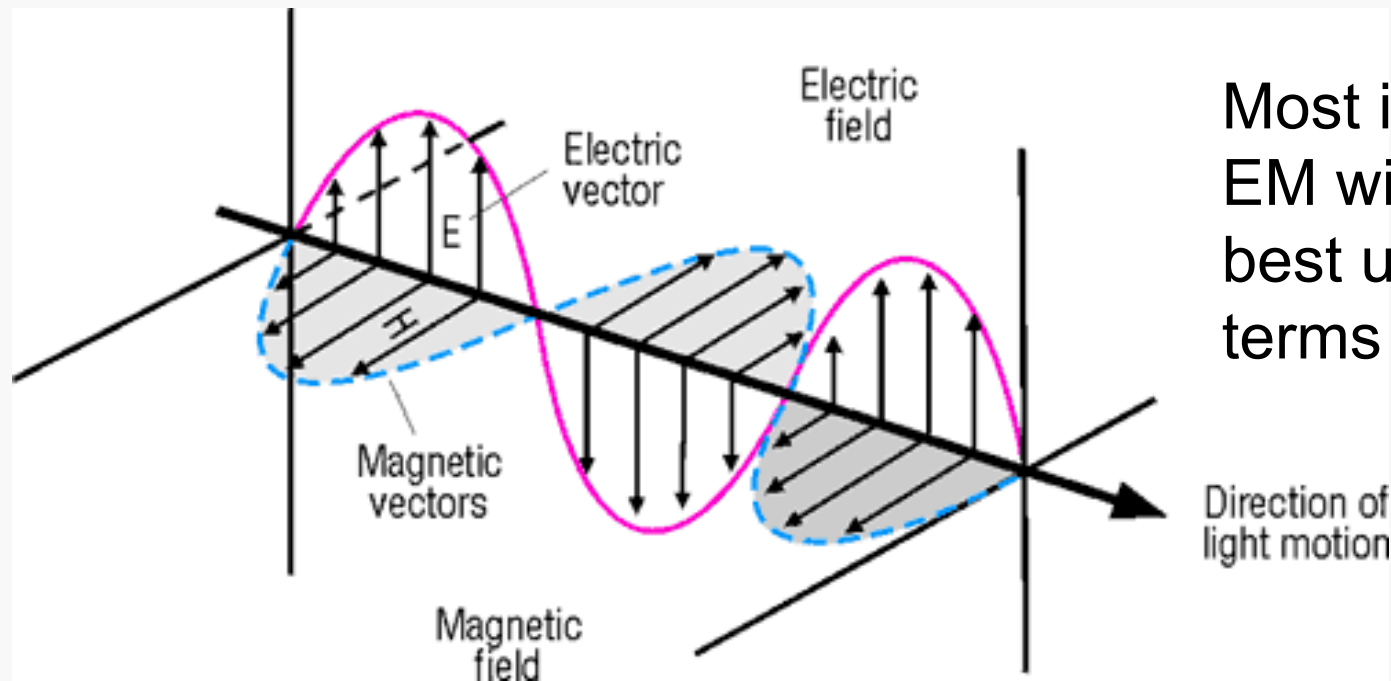


R O Y G B V

**Spectroscopy** = methods based on the interaction of electromagnetic radiation (EM) and matter

**Electromagnetic Radiation** = form of energy with both wave and particle properties

EM moves through space as a wave



Most interactions of EM with matter are best understood in terms of electric vector

## Relationship between various wave properties

$$\nu \lambda_i = \frac{C}{\eta_i}$$

Where  $\nu$  = frequency in cycles/s or Hz

$\lambda_i$  = wavelength in medium i

$\eta_i$  = refractive index of medium i

$C$  = speed of light in vacuum ( $2.99 \times 10^{10}$  cm/s)

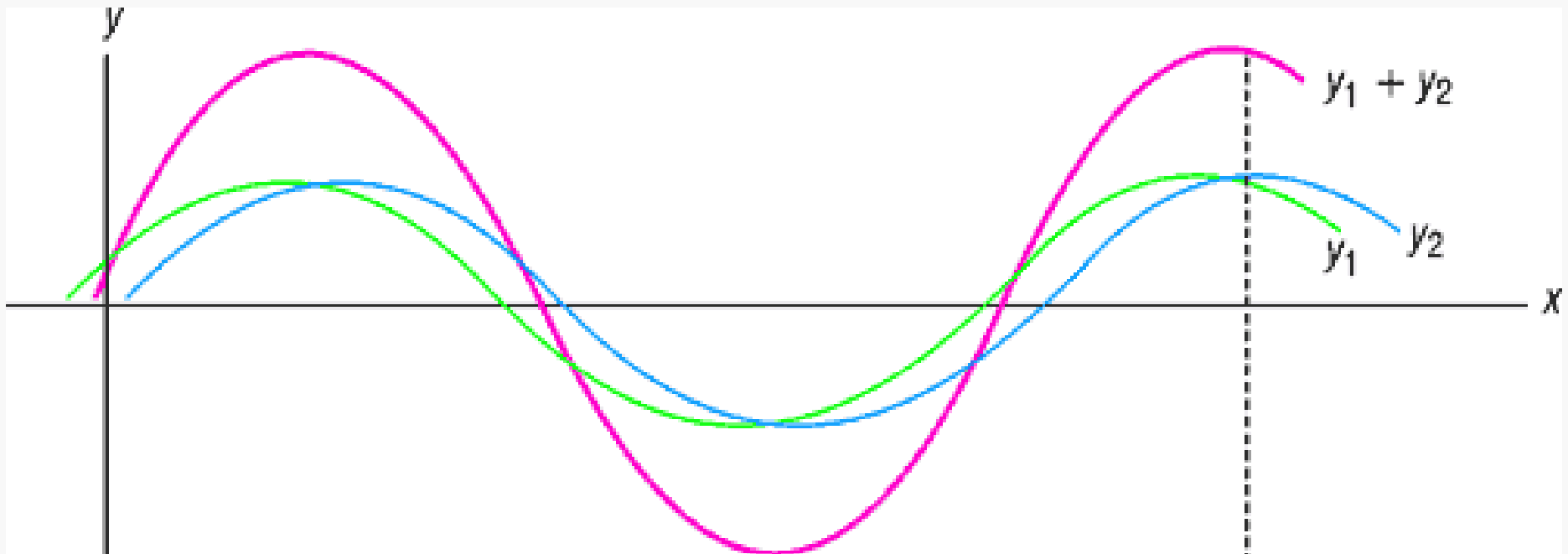
EM slows down in media other than vacuum because electric vector interacts with electric fields in the medium (matter) → this effect is greatest in solids & liquids, in gases (air) velocity similar to vacuum

## Wave Interaction - interaction between waves

- waves must have similar  $v$  but can be out of phase (i.e., they start in different places)

Principle of superposition = vectors add

- wave  $y_1 + y_2$  formed by adding  $y_1$  &  $y_2$  by vector addition



## Wave Equation

$$y = A \sin (\omega t + \alpha)$$

Where  $A$  = amplitude

$\omega$  = angular frequency

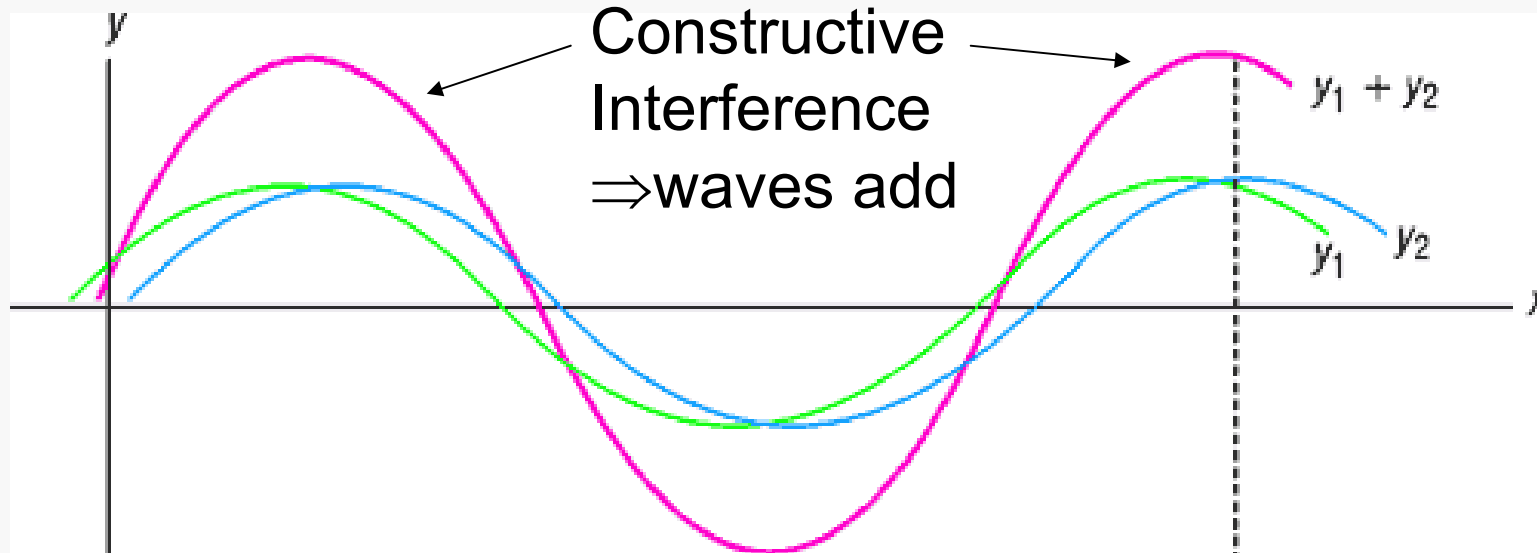
$\alpha$  = phase angle

$t$  = time

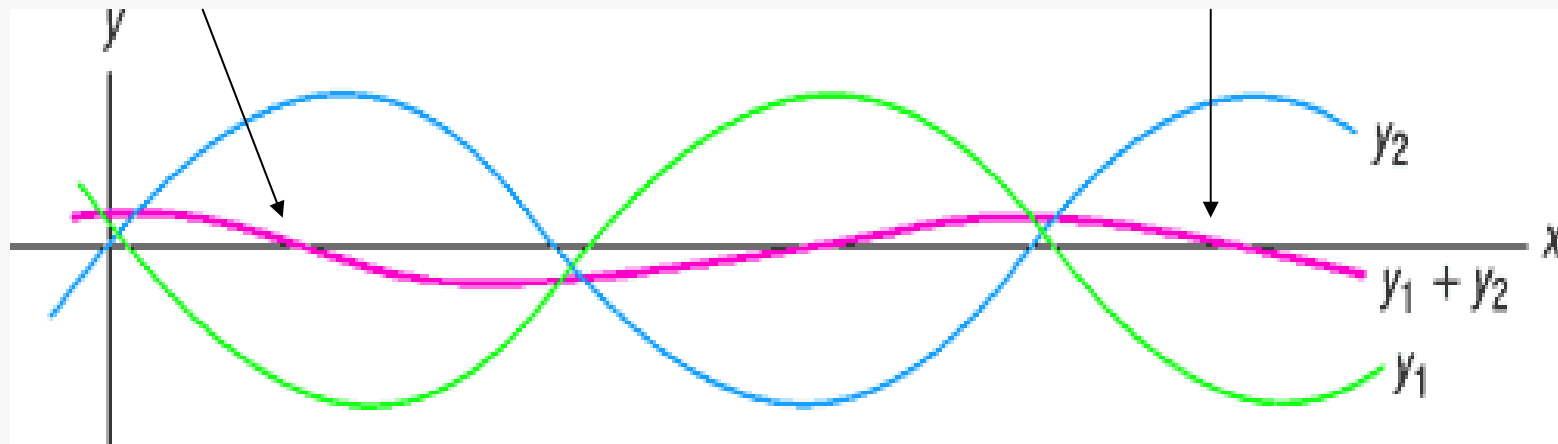
For a collection of waves the resulting position  $y$  at a given  $t$  can be calculated by

$$y = A_1 \sin (\omega_1 t + \alpha_1) + A_2 \sin (\omega_2 t + \alpha_2) + \dots$$

**Interference** - amplitude of the resulting wave depends on phase difference  $\alpha_1 - \alpha_2$

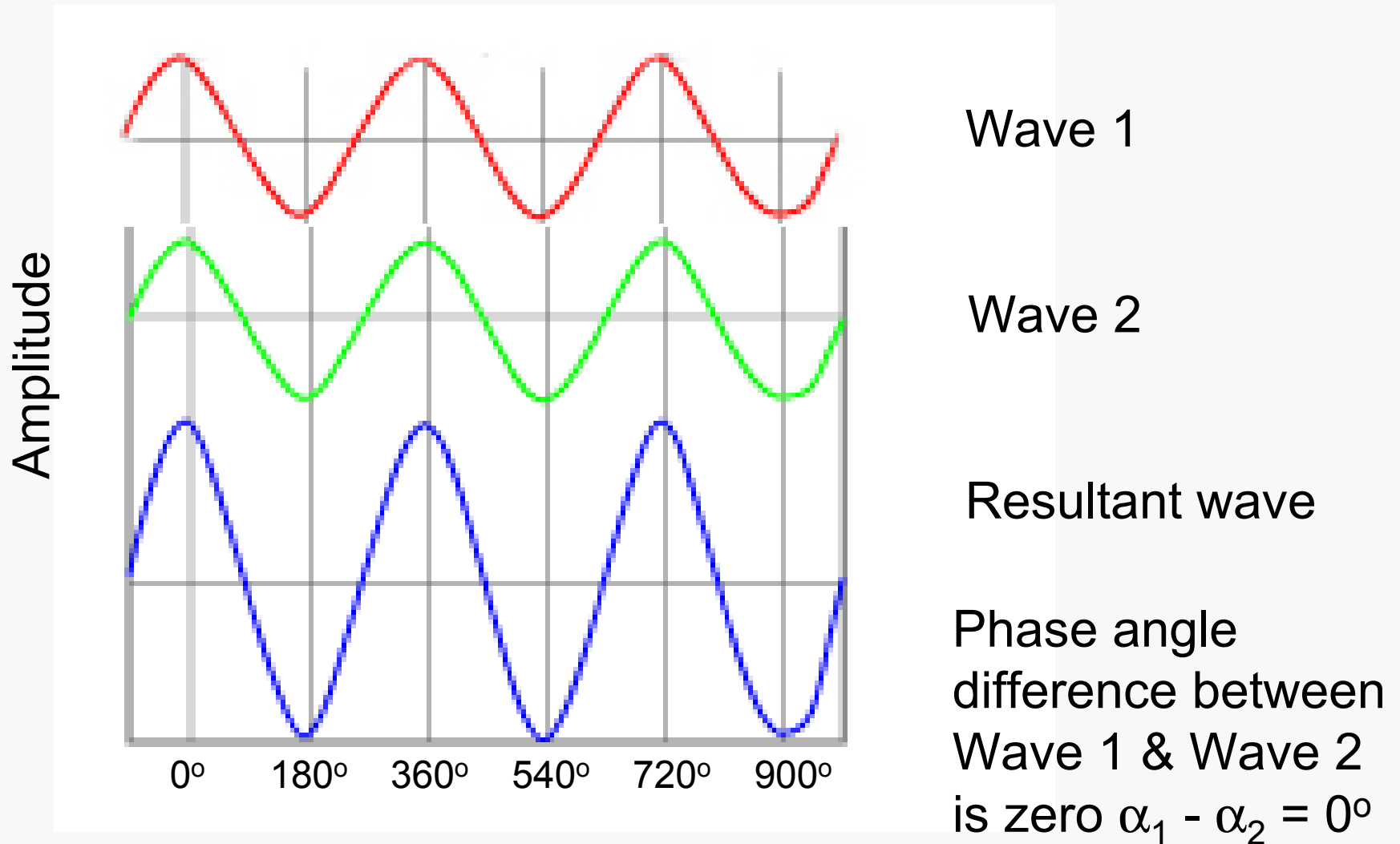


Destructive Interference  $\Rightarrow$  waves cancel

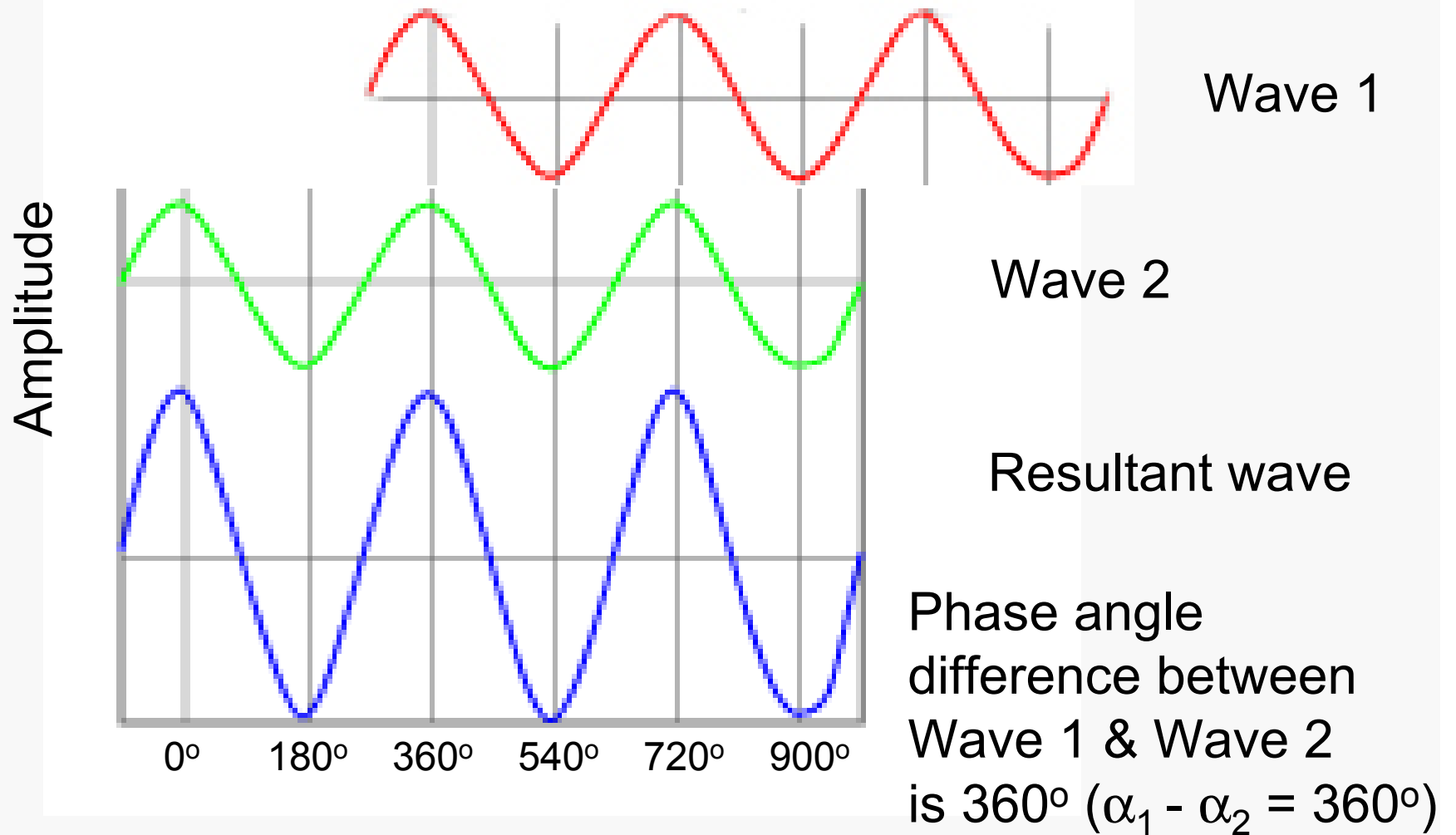




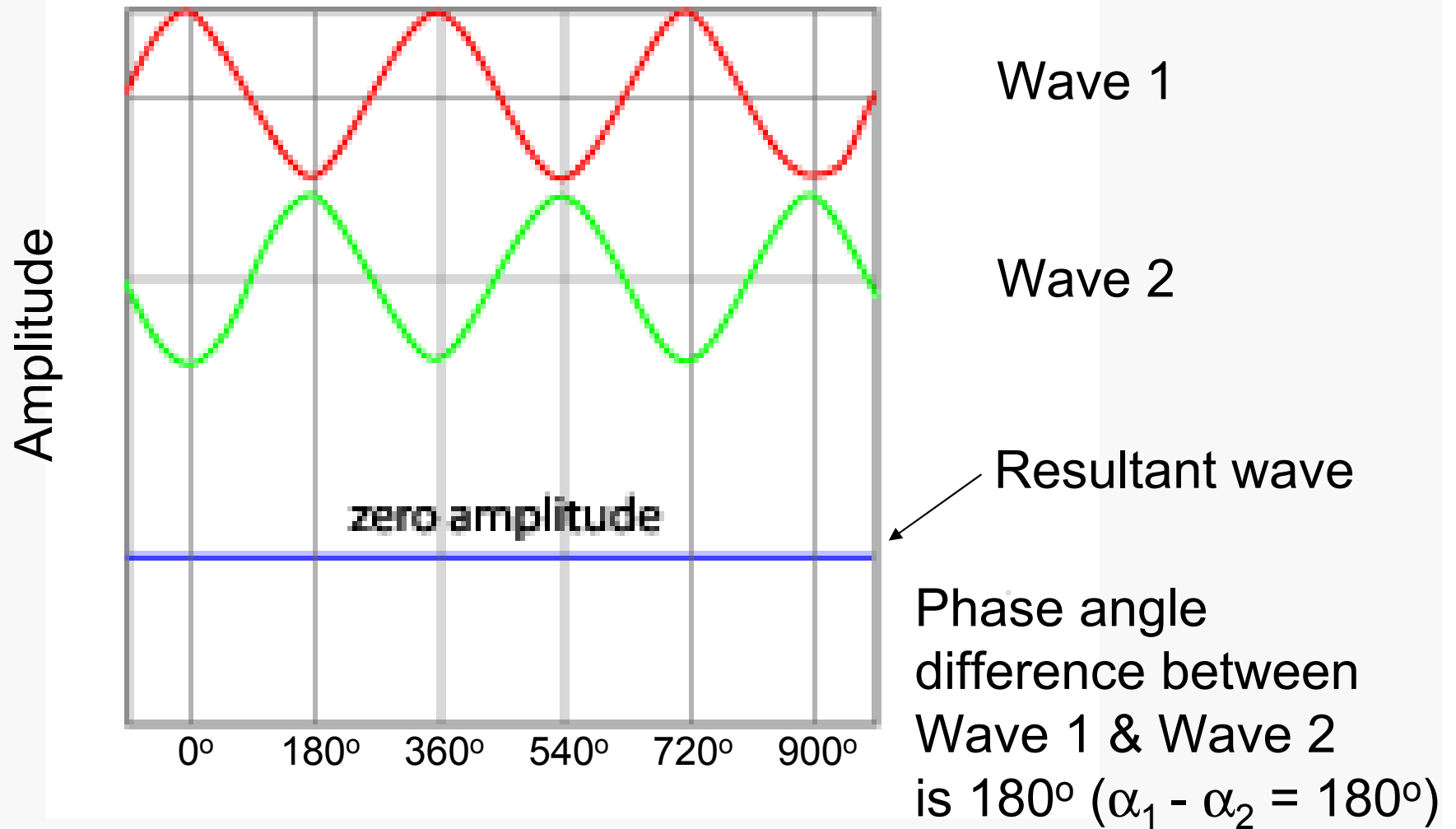
At  $\alpha_1 - \alpha_2 = 0^\circ$  adding of waves gives  
Maximum Constructive Interference



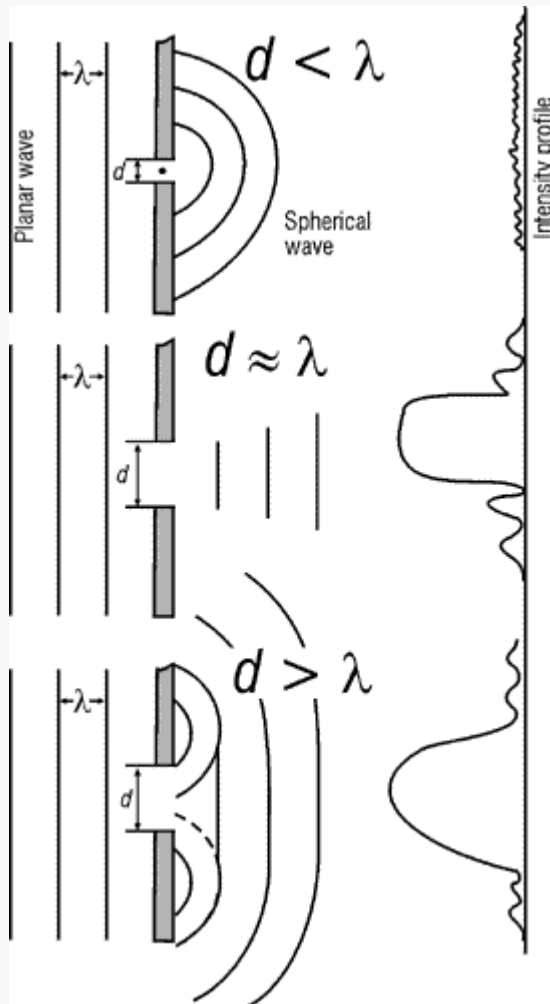
Also at  $\alpha_1 - \alpha_2 = 360^\circ$  adding of waves gives  
Maximum Constructive Interference



When  $\alpha_1 - \alpha_2 = 180^\circ$  or  $540^\circ$  adding of waves gives Maximum Destructive Interference



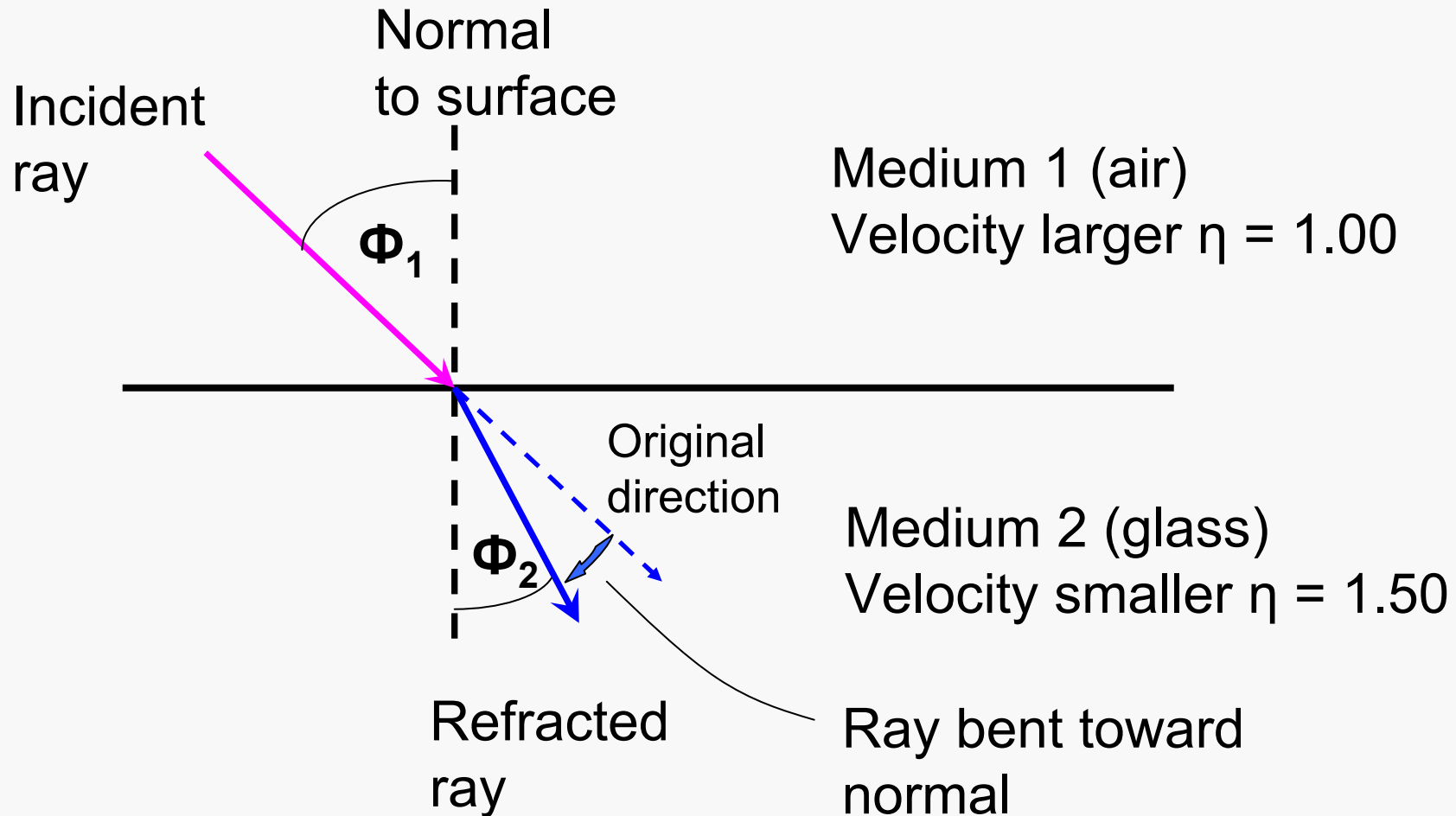
Diffraction = EM going past an edge or through a slit (2 edges) tends to spread



The combination of diffraction effects & interference effects are important in spectroscopy for

- 1) diffraction gratings
- 2) slit width considerations

**Refraction** = change in velocity of EM as it goes from one medium to another



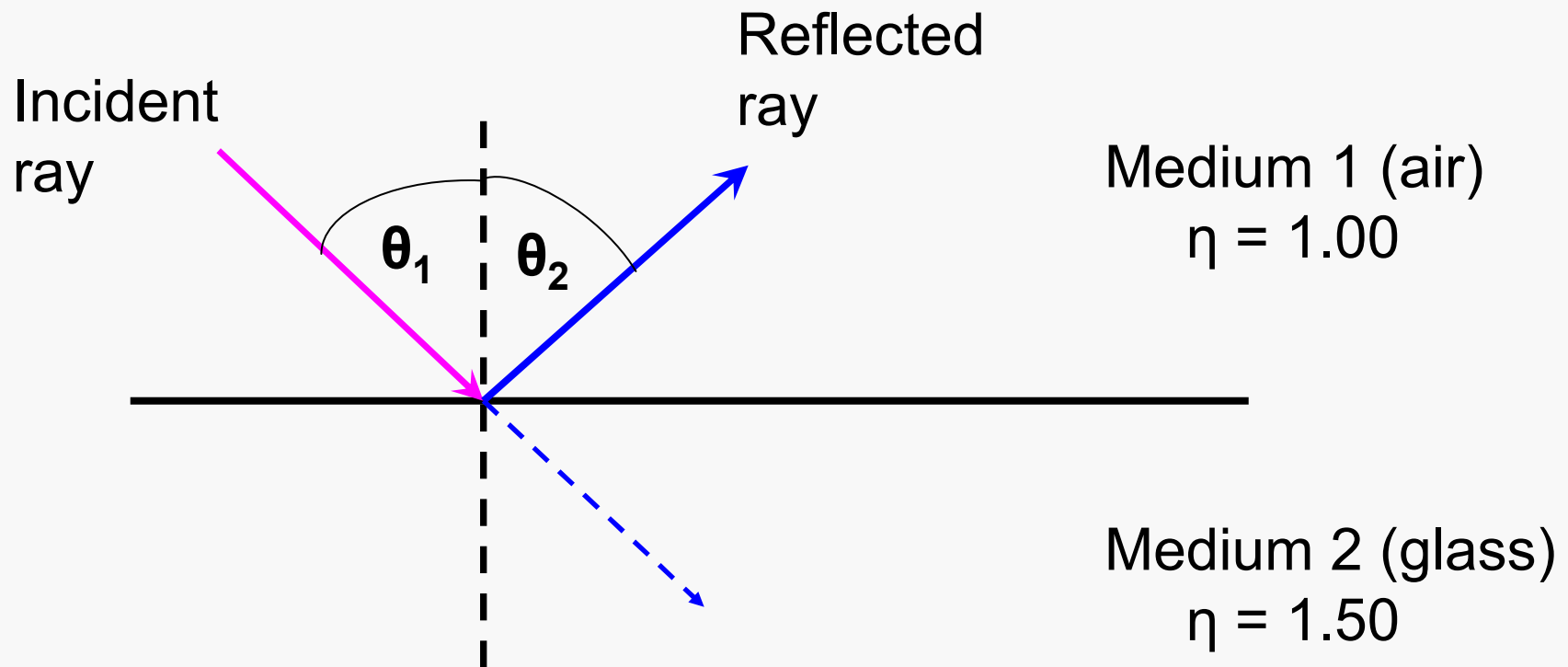
## Equation for Refraction

$$\frac{\sin \Phi_1}{\sin \Phi_2} = \frac{v_1}{v_2} = \frac{\eta_2}{\eta_1} = \eta_2 \quad \begin{array}{l} \text{if medium 1} \\ \text{is air } \eta_1 = 1.0 \end{array}$$

Magnitude of the direction change (i.e., size of the angle depends on wavelength (shown in equation as  $v$ ) this is how a prism works

Direction of bending depends on relative values of  $\eta$  for each medium. Going from low  $\eta$  to higher, the ray bends toward the normal. Going from higher  $\eta$  to lower the ray bends away from the normal.

**Reflection** = EM strikes a boundary between two media differing in  $\eta$  and bounces back



Specular reflection = situation where angle of incidence ( $\theta_i$ ) equals angle of reflection ( $\theta_r$ )

$$\text{Reflectance} = R = \frac{I_r}{I_i} = \frac{(\eta_2 - \eta_1)^2}{(\eta_2 + \eta_1)^2}$$

Where  $I_i$  and  $I_r$  = incident & reflected intensity

For radiation going from air ( $\eta = 1.00$ ) to glass ( $\eta = 1.50$ ) as shown in previous slide

$$R = 0.04 = 4 \%$$

Many surfaces at 4 % each (i.e., many lenses) can cause serious light losses in a spectrometer. This generates **stray radiation** or **stray light**.